A banner for the Mars Exploration Program Analysis Group (MEPAG). The background is a dark blue, textured surface, possibly a Mars rock or soil, with a large, reddish-orange sphere of Mars on the right side. The text is white and bold.

Mars Exploration Program Analysis Group (MEPAG)

chartered by NASA HQ to assist in planning the scientific exploration of Mars


A dark, grainy image of the Mars surface, showing a horizon line and several small white stars in the sky.

Mars: Current State of Knowledge and Why Mars Remains a Compelling Objective

Jack Mustard, *MEPAG* Chair,

On behalf of the Mars Exploration Program
Analysis Group

Sept. 9, 2009

A banner for the Mars Exploration Program Analysis Group (MEPAG). The background is a dark blue, textured surface, possibly representing a Mars crater or a similar geological feature. On the right side, a portion of the reddish-orange planet Mars is visible. The text is white and positioned on the left side of the banner.

Mars Exploration Program Analysis Group (MEPAG)

chartered by NASA HQ to assist in planning the scientific exploration of Mars

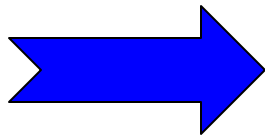
A large, high-resolution image of the Mars surface, showing a vast, flat, reddish-orange landscape with subtle variations in color and texture. The image is centered and occupies most of the slide area.

What Were Our Goals for the Past Decade?

MEPAG's Goals and Strategies, 2001-2011

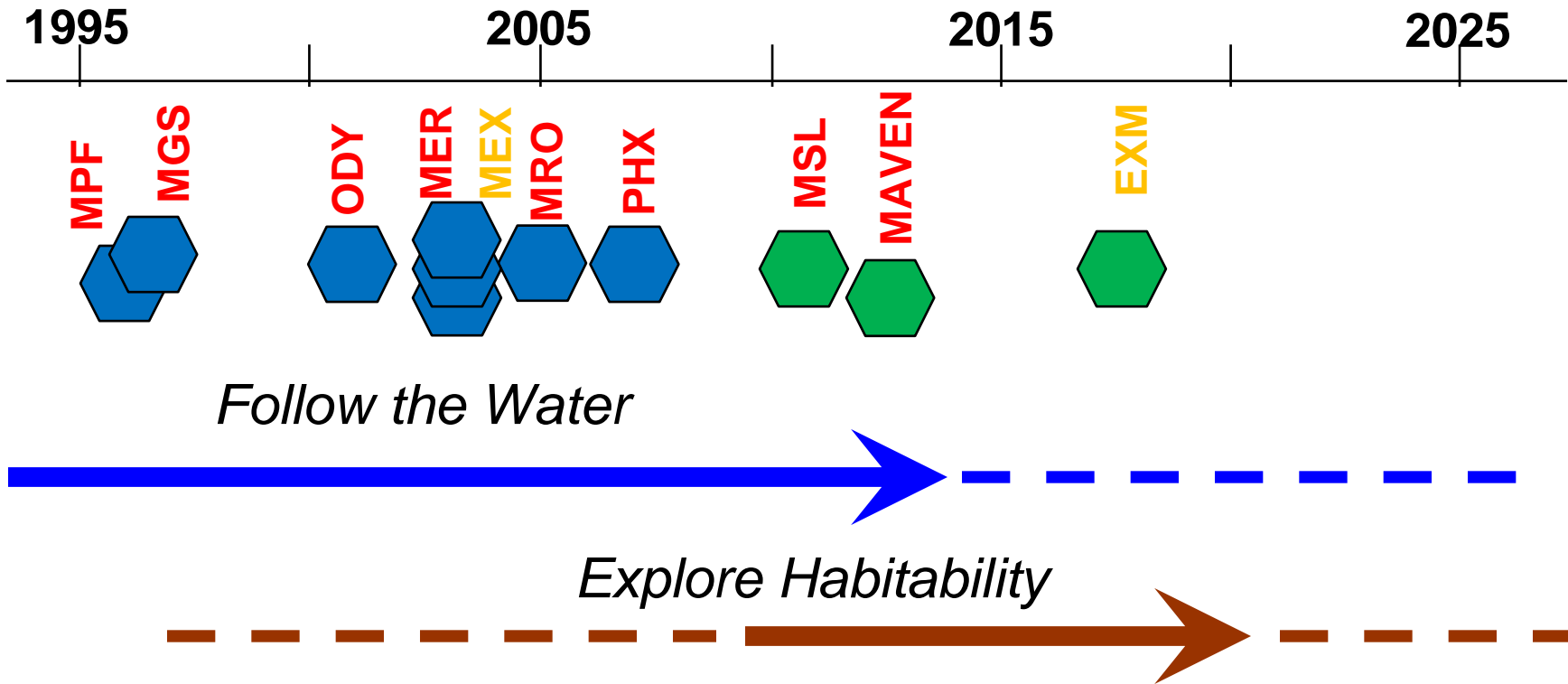
- I. Determine if life ever arose on Mars
- II. Understand the processes and history of climate on Mars
- III. Determine the evolution of the surface and interior of Mars
- IV. Prepare for eventual human exploration

2001 Strategy
Follow the Water



2005 Strategy
Explore Habitability

Missions In Progress to Address Goals



Missions Legend

-  **Successfully Flown**
-  **In Development**

Mars Exploration Program Analysis Group (MEPAG)

chartered by NASA HQ to assist in planning the scientific exploration of Mars

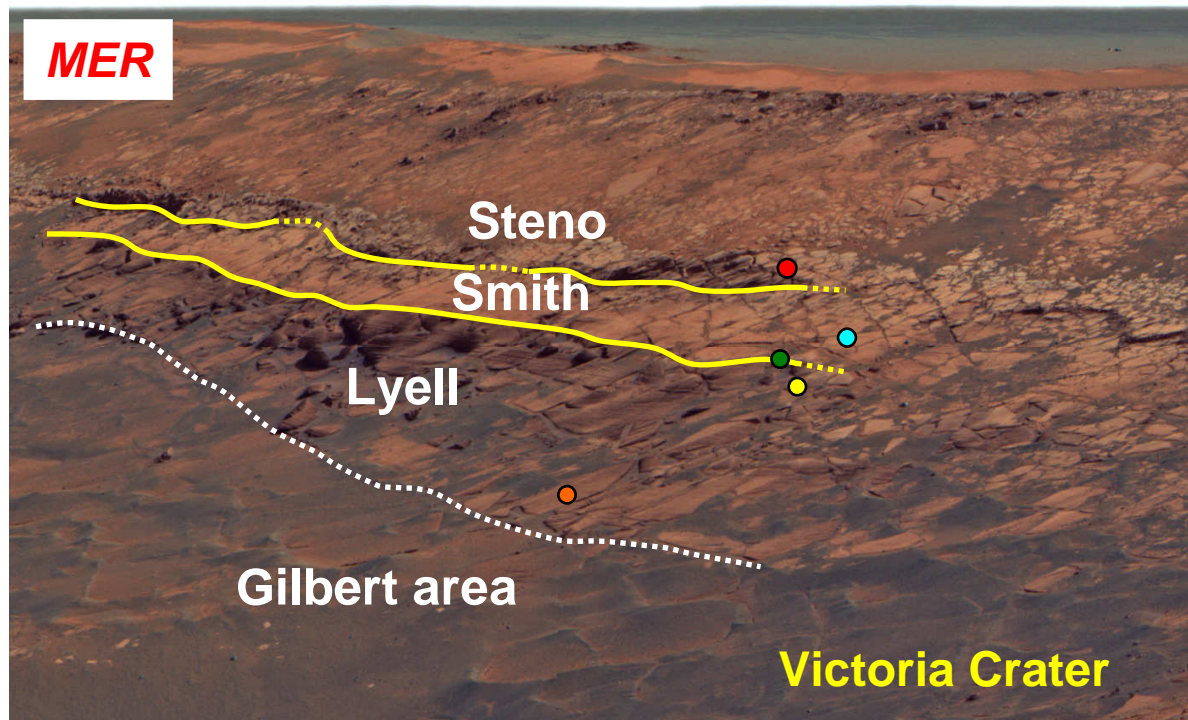
What Did We Learn?

Last Decade Discoveries: Diverse Planet with Complex History

- ❑ We have made significant advances in understanding the processes and history of climate, as well as understanding the evolution of the surface (Goals II, III).
 - Mars has areas with diverse mineralogy, including alteration by water, with a change in mineralogy over time [MGS, ODY, MER, MEX, MRO]
 - In situ confirmation of Wet (Warm?) Climate in the past [MER]
 - Pervasive water ice in globally distributed, near-surface reservoirs [ODY, MRO, MEX, PHX]
 - Sources, phase changes, and transport of volatiles (H₂O, CO₂) are known & some are quantified [MGS, MEX, MRO, PHX]
 - Increasing evidence for geologically recent climate change: stratified layers in ice and in rock [MGS, ODY, MEX, MRO]
 - Dynamic change occurring even today: landslides, new gullies, new impact craters, changing CO₂ ice cover [MGS, ODY, MEX, MRO]
 - Presence of methane indicative of active chemical processes either biogenic or abiotic [MEX and ground-based]
- ❑ Based on much of the above, the perception of *Potential for past Life* has increased, and *Modern Life* may still be possible. (Goal I)

Last Decade Discoveries: Diversity of Environments

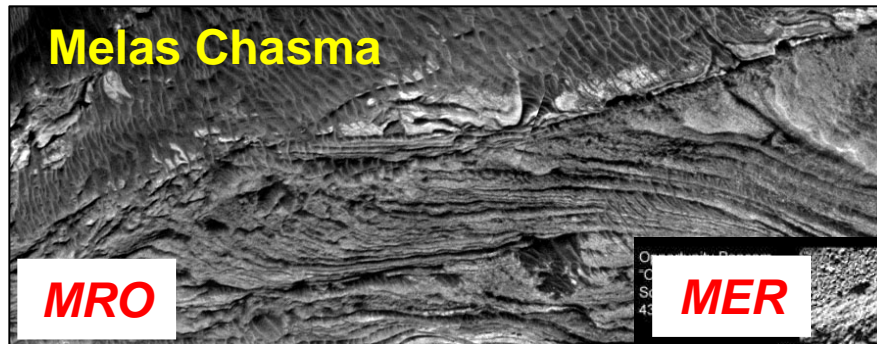
- ❑ Chemistry and morphology indicate changing environments throughout geologic history
 - Acidic waters at Meridiani
 - Basic waters at Phoenix landing site
 - Mineralogy: clays to sulfates to oxides



Hesperian subsurface water, diagenesis

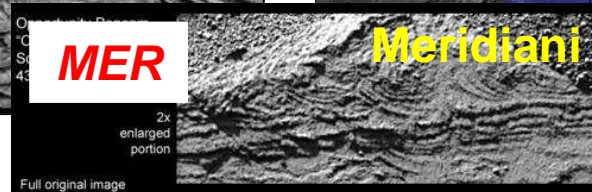


Past Decadal Results: Wide variety of sedimentary deposits

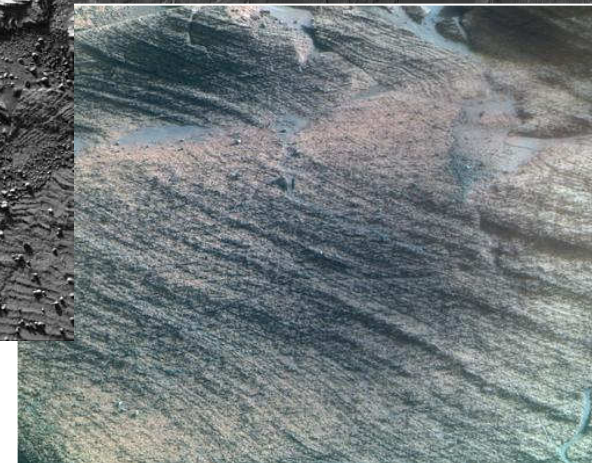
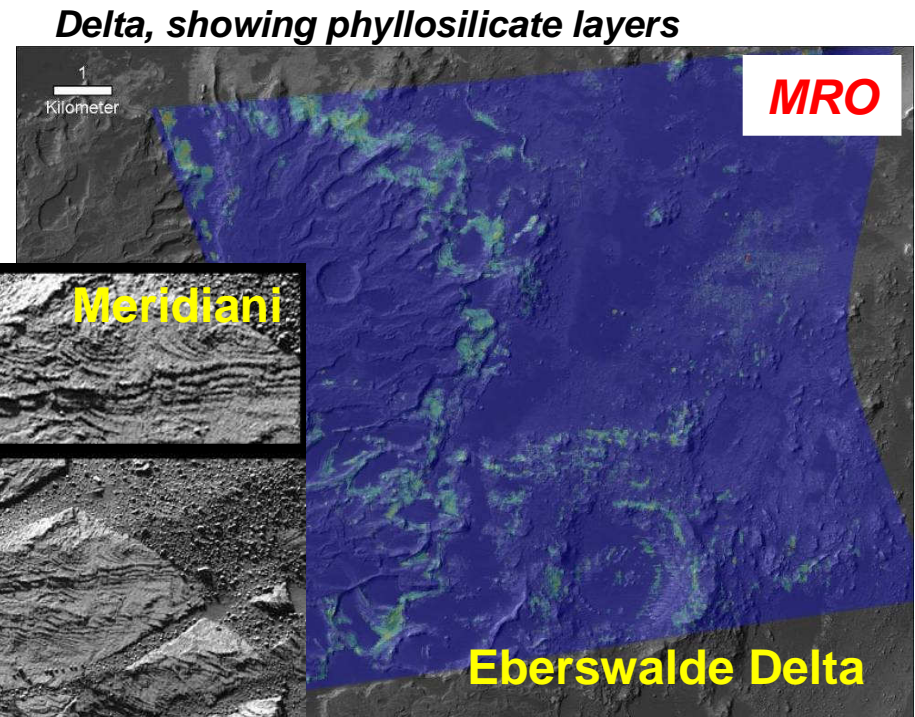


Large-scale sedimentary structures

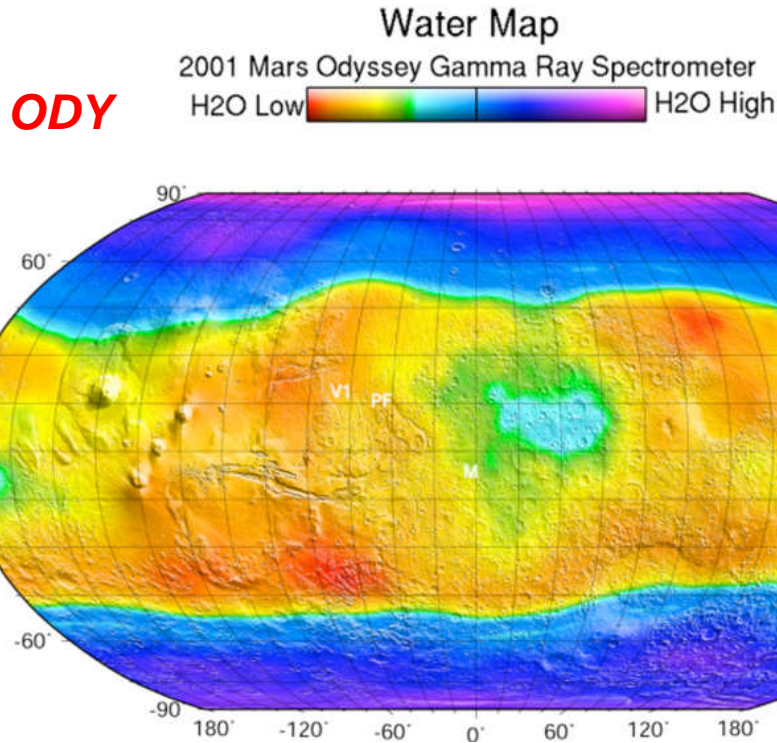
- **Depositional processes created a sedimentary record**
 - Developed in topographically low areas
 - Spectacular stratification at multiple scales
 - Evidence of persistent standing water, lakes
 - Sediments systematically change in character with time
 - Multiple facies recognized



Fine-scale sedimentary structures



Past Decadal Results: Distribution of Modern Water



Global Near-Surface Reservoirs of Water

Gamma Ray Spectrometer

- Global hydrogen abundance and equivalent H₂O
- Ground ice to +/-60° in high abundance

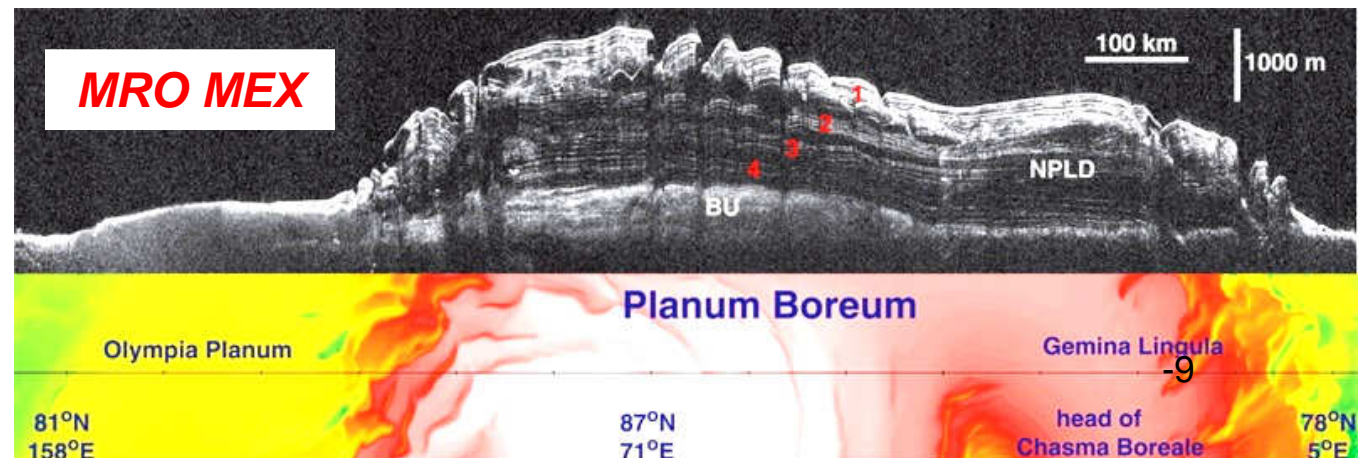


Phoenix results

PHX

SHARAD and MARSIS

- Nearly pure water ice
- Distinct layering
- No deflection of crust
- Ice-cored lobate debris aprons in mid-latitudes

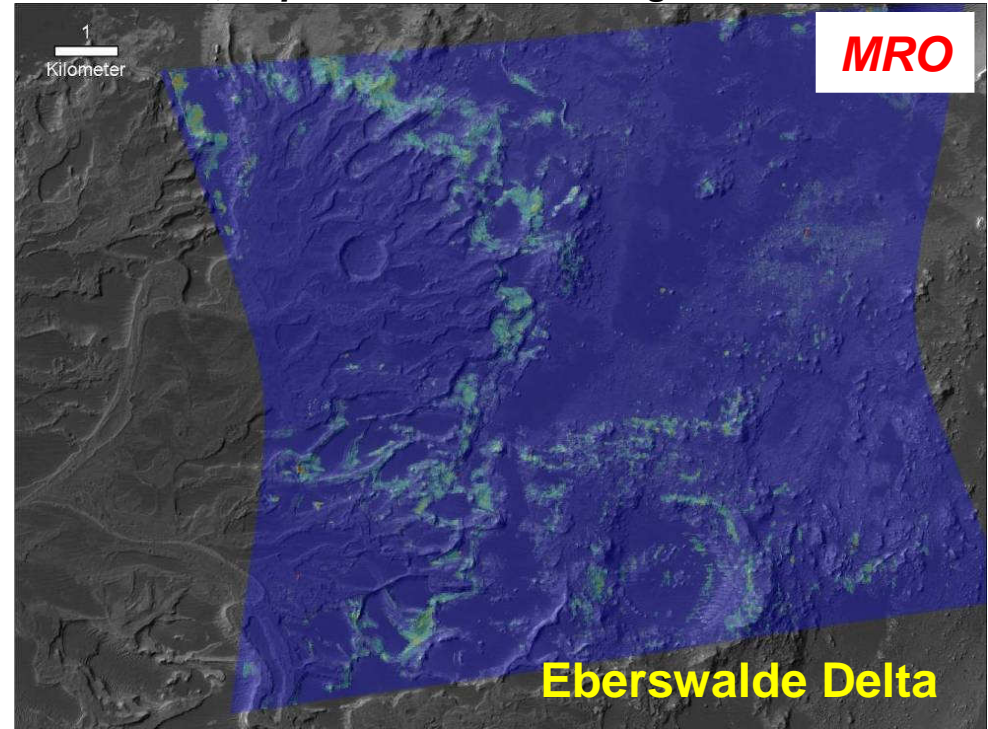


Past Decadal Results: Ancient Mars Was Wet (Episodically?)

Channels formed by rainfall runoff

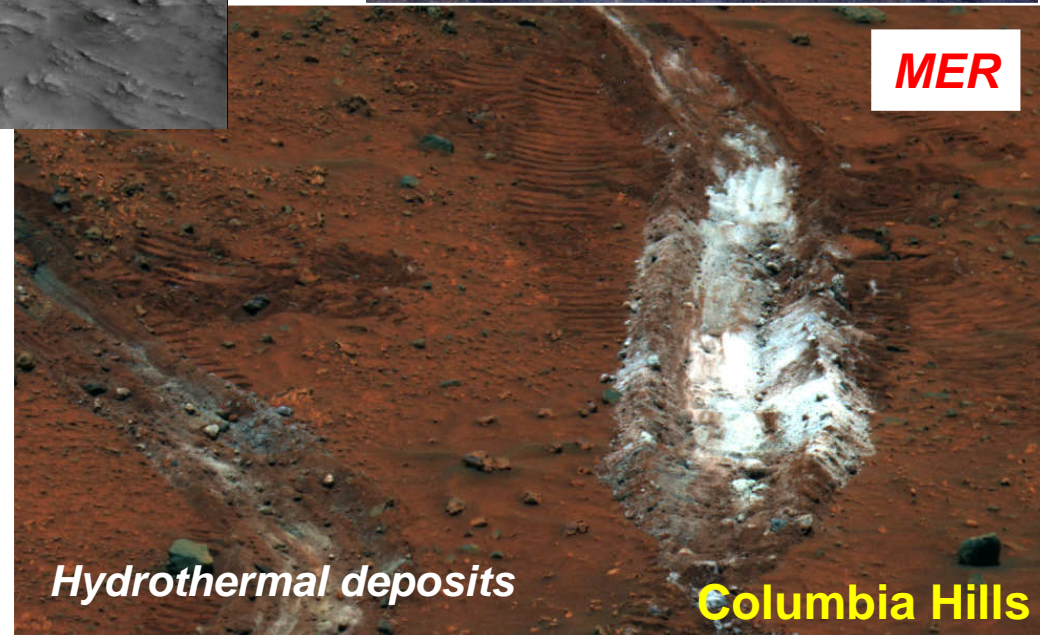
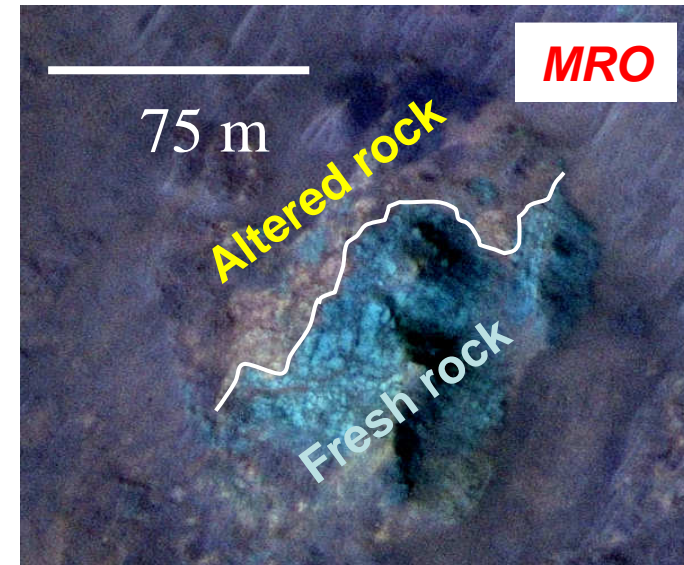
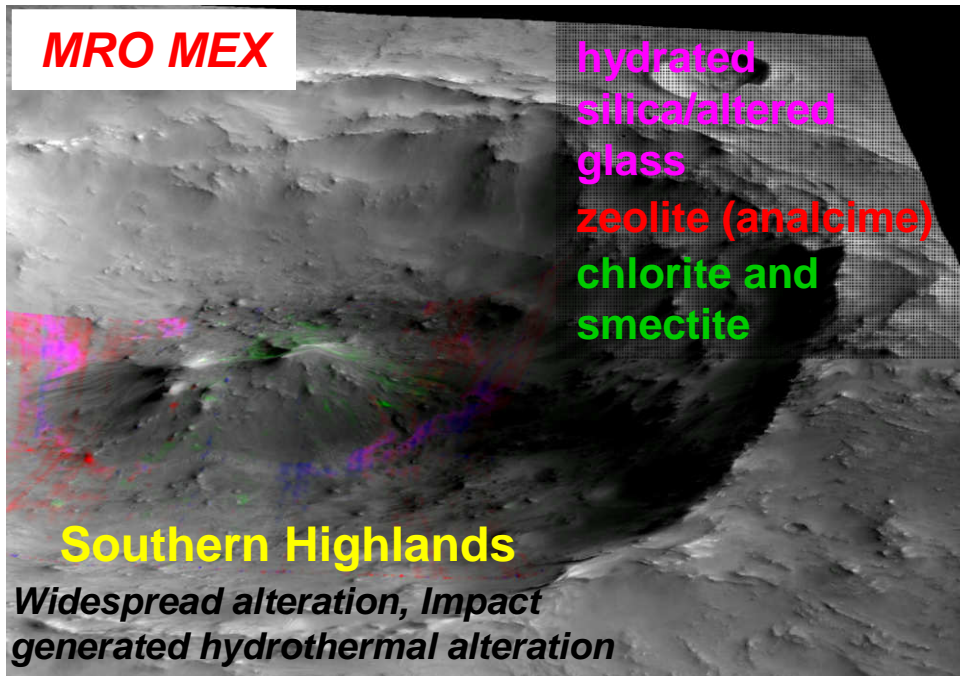


Delta, deposition into standing water

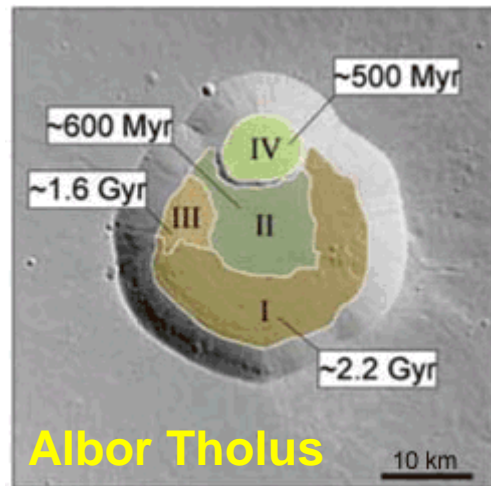
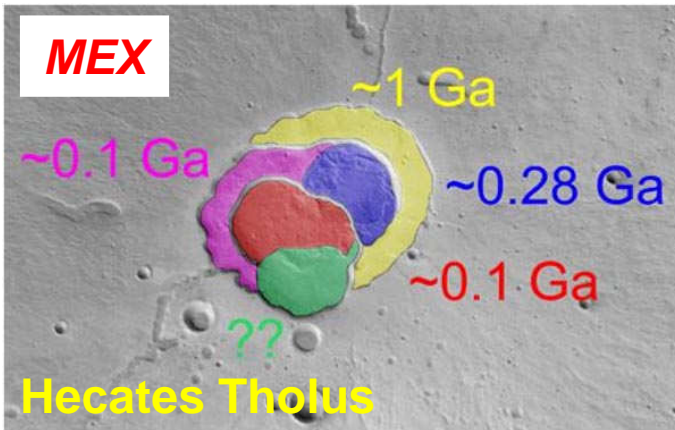


- ❑ Ancient features indicate water present at the surface
 - Evidence of persistent standing water, lakes
 - Evidence of rainfall, valley networks
 - Lake overflow features

Past Decadal Results: Evidence for Water/Rock Interaction

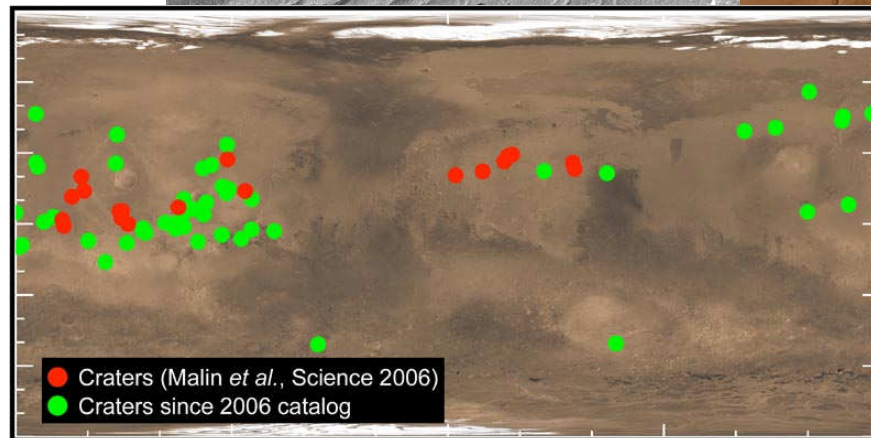
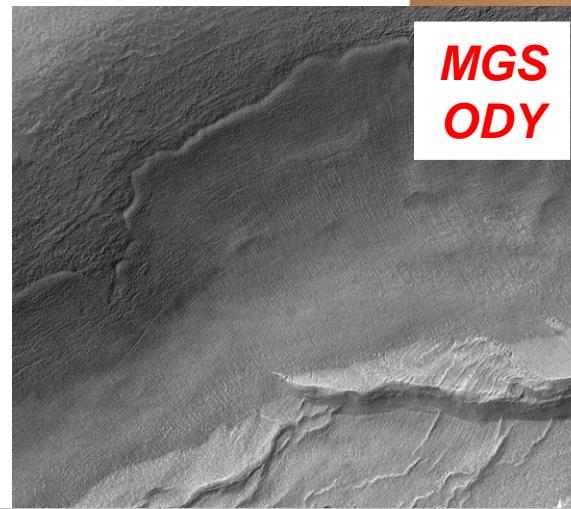


Past Decadal Results: Mars Still Active Today



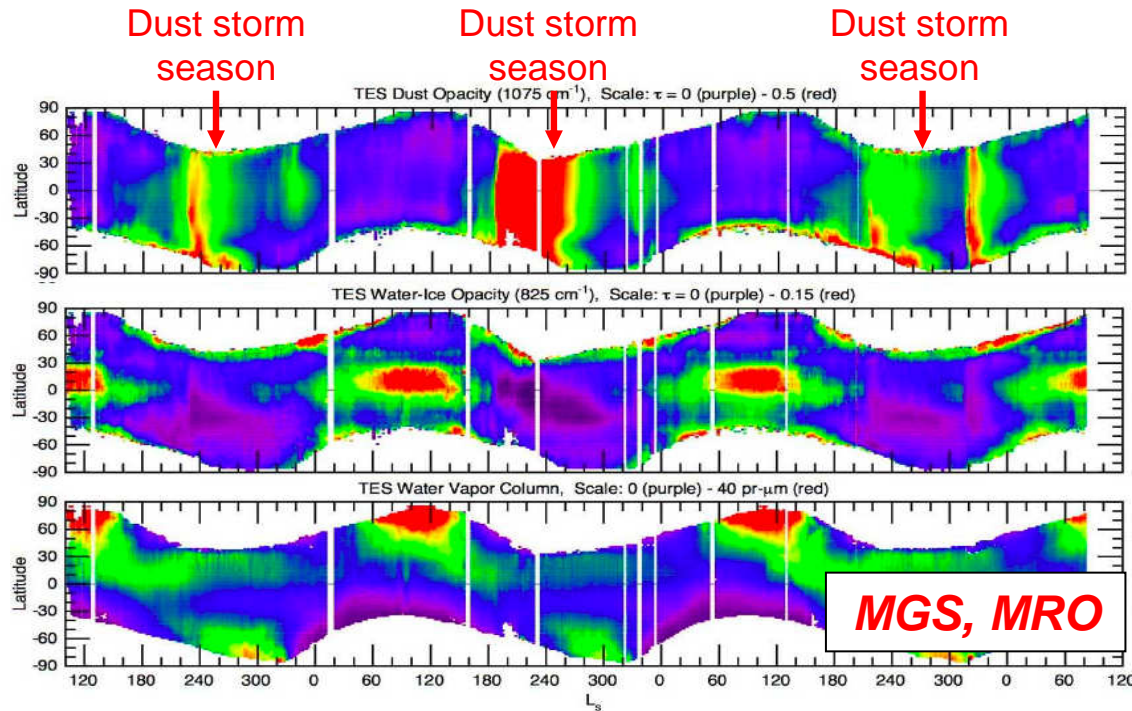
Volcanic activity spans most or all of martian geologic history

Mid-latitude mantles and gullies

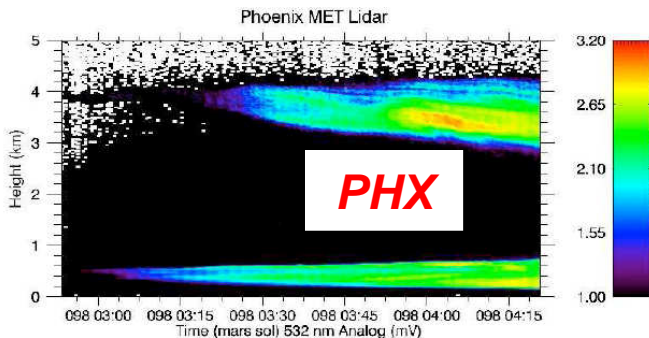


New Impact Craters

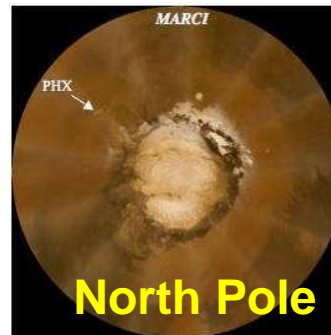
Past Decadal Results: Atmosphere and Climate Results



Understand how the atmosphere works



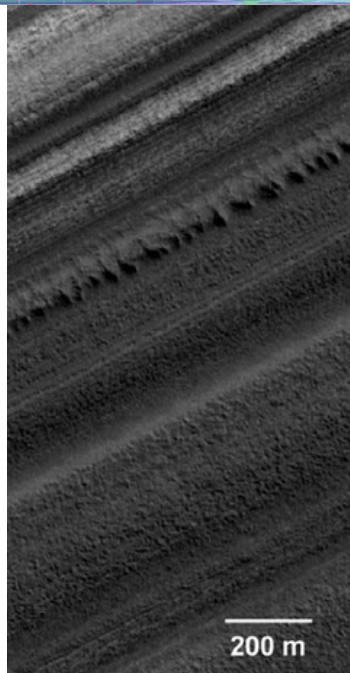
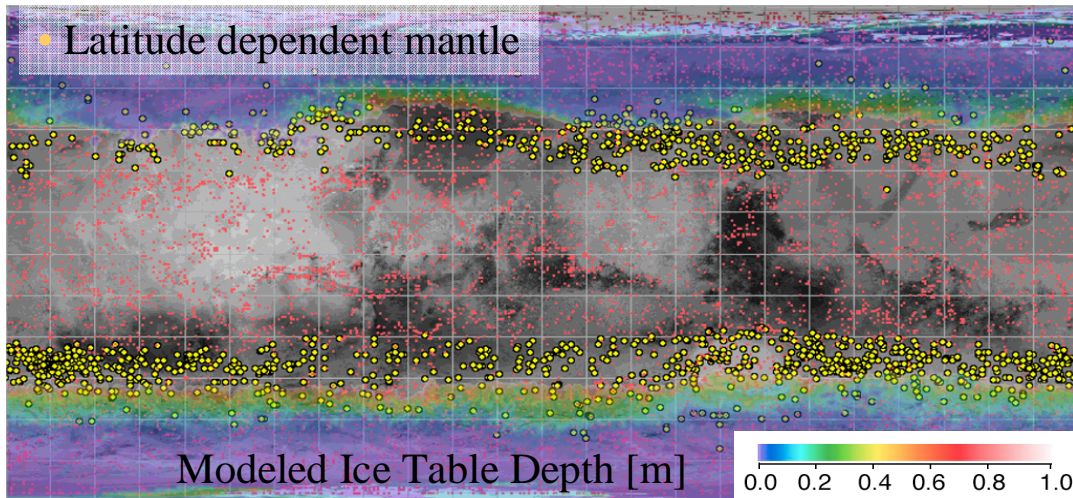
Cloud, fog and storm dynamics



MEX, MRO

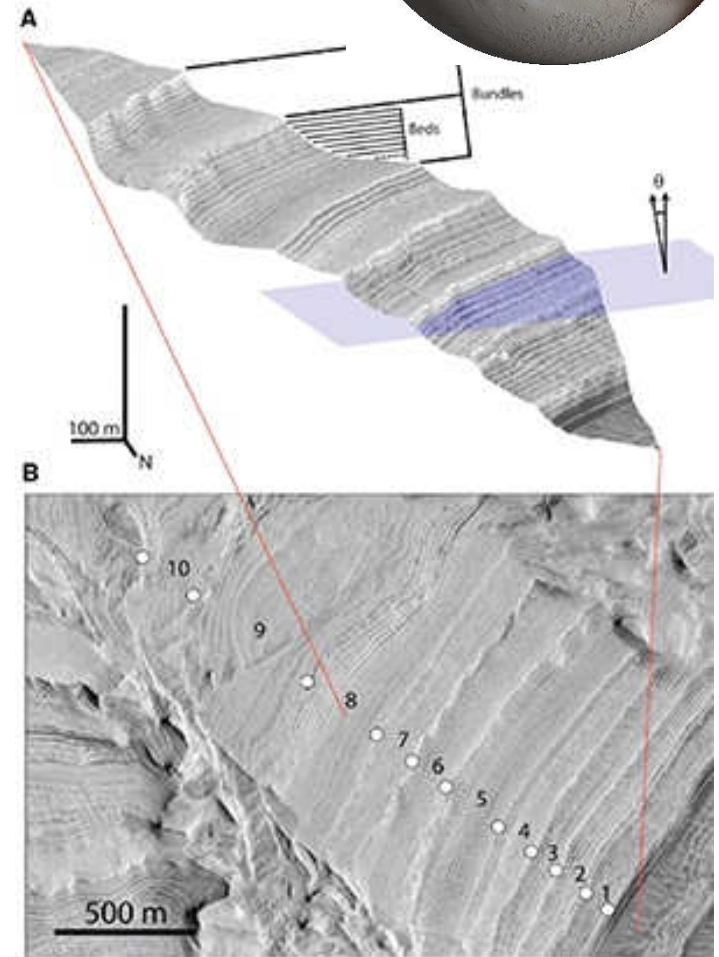
- Climate change -- Past, recent and past: Understanding the process
 - Early wet (warm?) Mars (Noachian) has evolved to cold, dry Mars (Hesperian +)
 - Periodic change in last several million years
- Recent multi-year record of CO₂/water/dust; atmospheric dynamics [MGS, ODY, MEX, MRO]
 - Seasonal cycles and interannual variability
- SO₂, Argon, CH₄, CO, etc.: Tracers of transport, chemistry, and surface-atmosphere interactions

Past Decadal Results: Periodic Climate Change



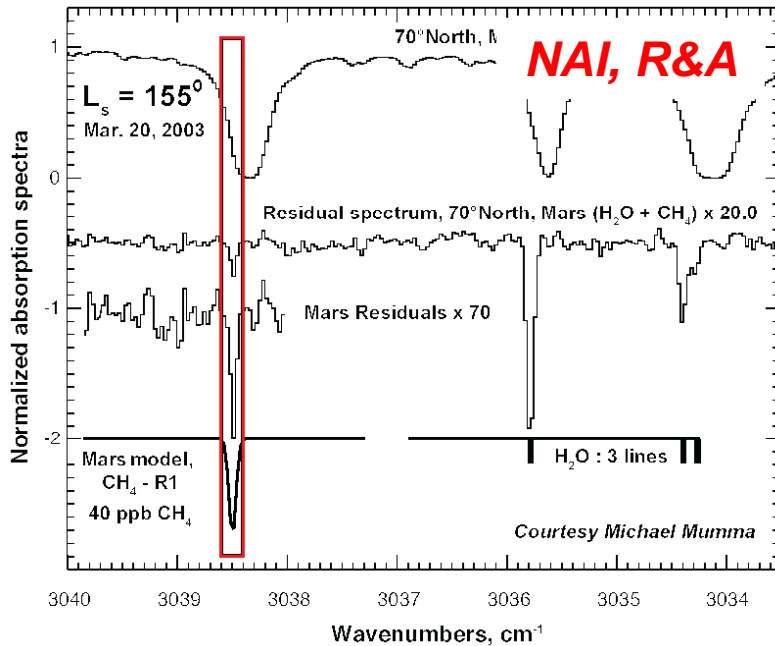
- Volatile-rich, latitude dependent deposits (mantle, glaciers, gullies, viscous flow) coupled to orbitally-forced climate change
- Periodicity of layering in the north polar cap deposits as well as sedimentary deposits

MGS, ODY, MEX MRO

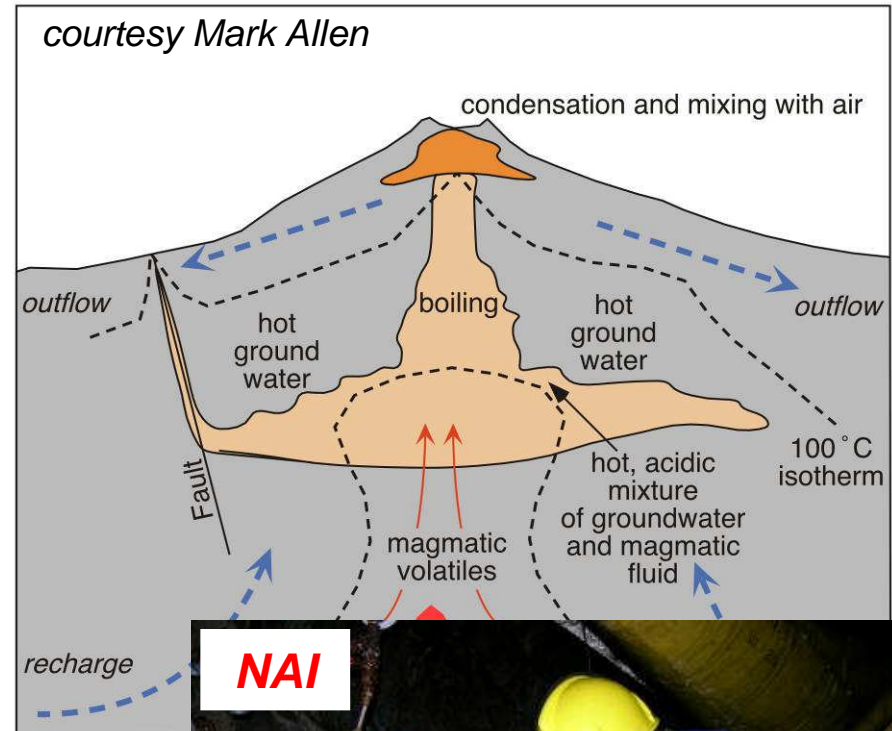


Past Decadal Results: Modern Methane

Courtesy Mike Mumma



Courtesy Michael Mumma



courtesy Lisa Pratt

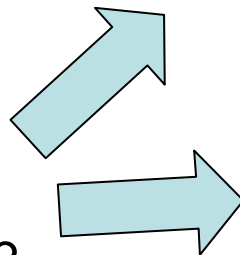
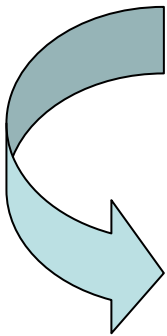
Detection of Methane on Mars

MEX NAI R&A

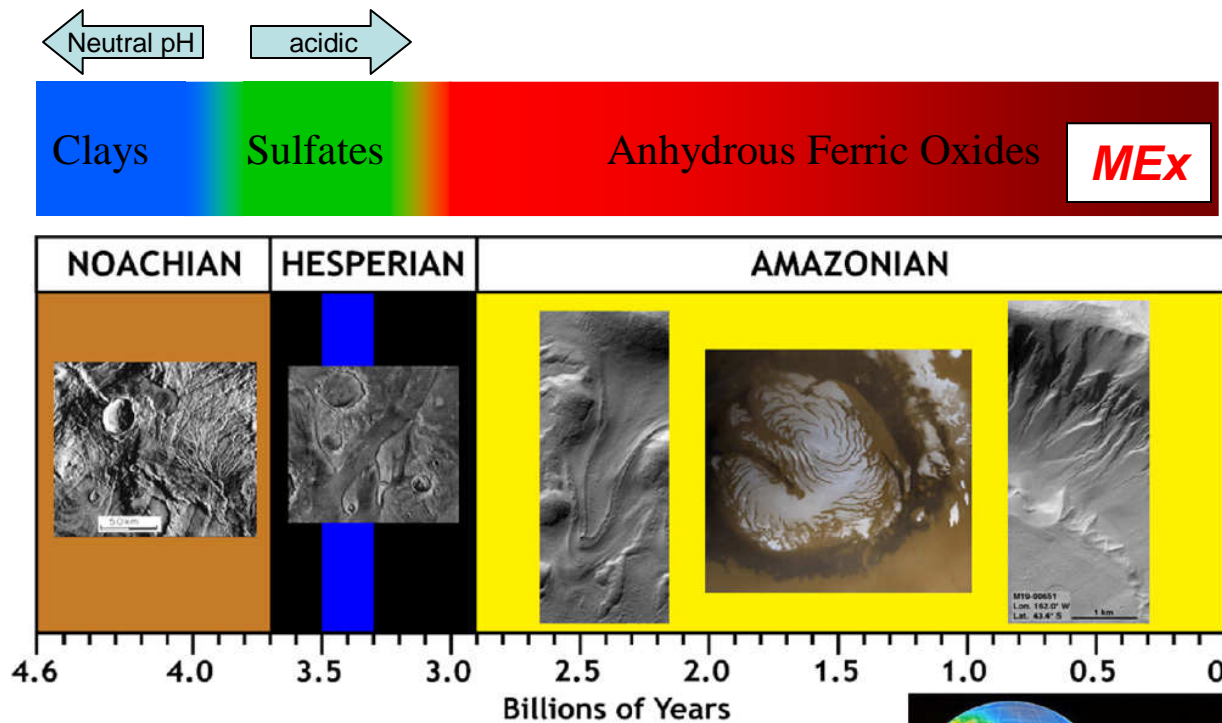
Abiotic?

Biotic?

Evidence of an active subsurface?



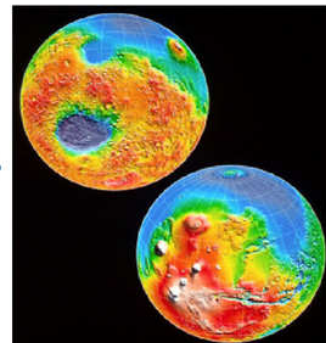
Past Decadal Results: Mars Planetary Evolution



- Heavy impact bombardment.
- Valley networks.
- "Warm/Wet" early Mars?

- Volcanism.
- Outflow channels.
- Oceans?
- South circumpolar deposits.

- Low impact rates.
- Tharsis volcanism continues.
- Outflow channels continue.
- Late-stage polar caps.
- "Cold/Dry" late Mars.



All Missions

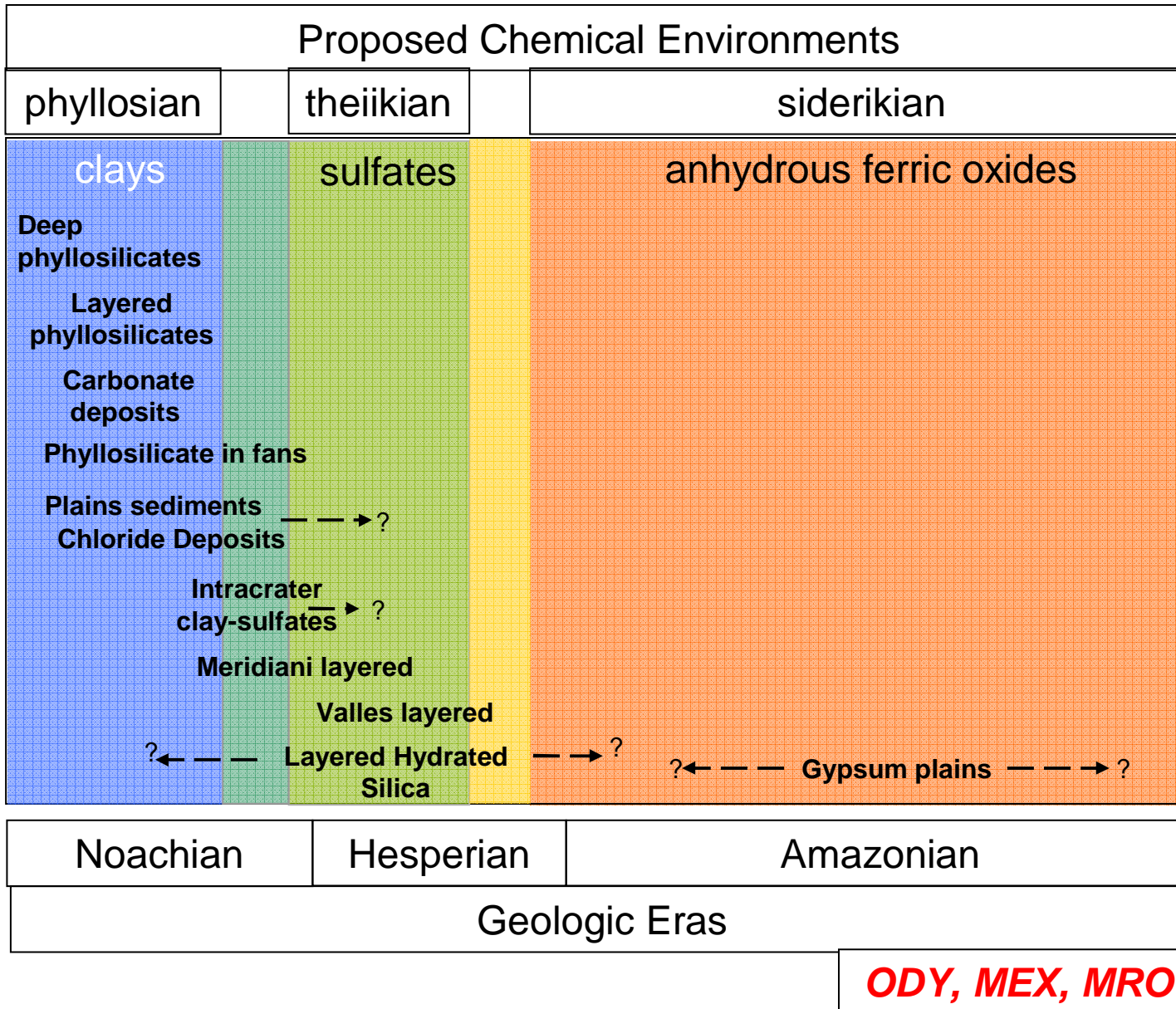
□ Hydrus Mineralogy Changed Over Time

- Phyllosilicate minerals (smectite clay, chlorite, kaolinite...) formed early
- Evaporates dominated by sulfate formed later with opal/hydrated silica
- Few hydrated mineral deposits since

□ Evolution of Aqueous, Fluvial and Glacial, Morphology with Time

- Valley networks, lake systems
- Gullies
- Viscous flow, glaciers, latitude dependant mantle

Past Decadal Results: Mars Planetary Evolution

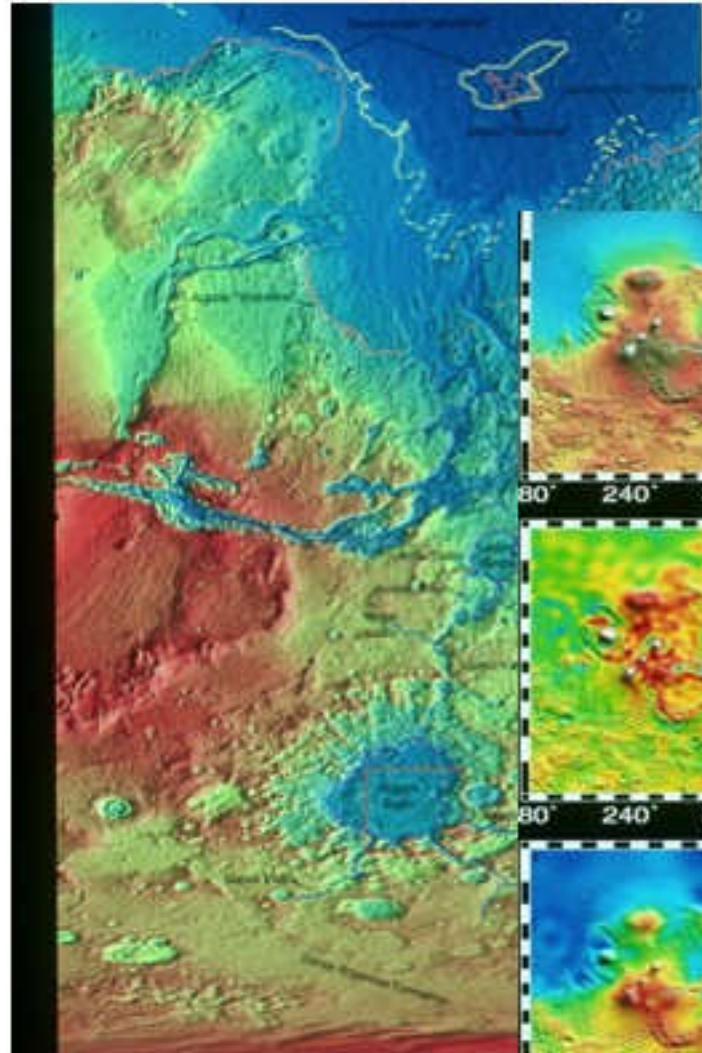


Coupled mineralogy and morphology define aqueous environments

Their character has evolved indicating changing environments

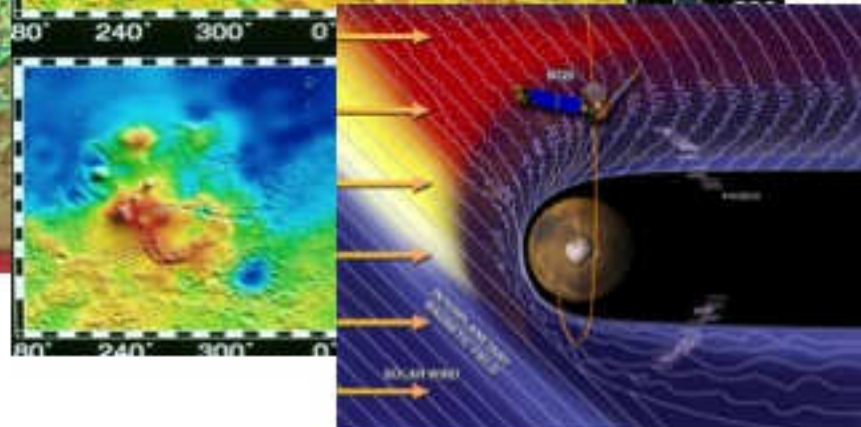
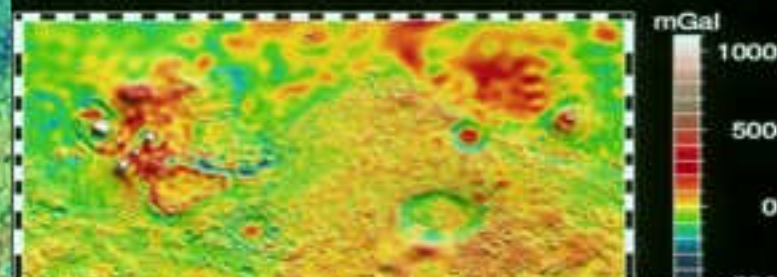
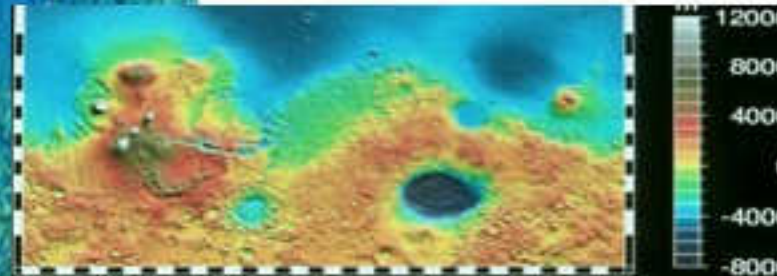
Data support the hypotheses but indicate greater complexity in local environments

Past Decadal Results: Crustal Structure and History



MGS MRO

- Best (30 km res) topographic model
 - Pole to pole slope controlling water transport, including northern depression



- Thick crust with near-isostatic compensation in rougher south
- Range of uncompensated gravity anomalies in smoother north
- Center of figure offset by nearly 3 km, indicating north pole ~6 km lower than south pole
- Evidence of early molten interior

Past Decadal Results:

Goal IV Prepare for Eventual Human Exploration

□ Following the water is a key first step in the preparation for human presence on Mars

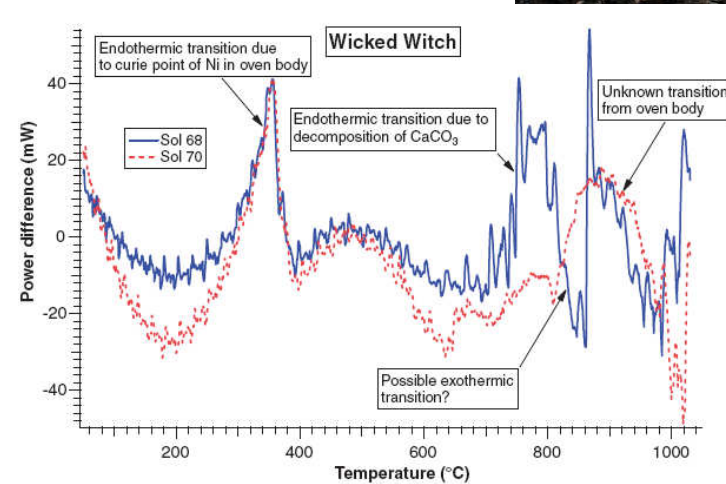
- Ice table at the depth, location, and concentration predicted by orbital data and theory


□ Phoenix instrumentation designed for environmental characterization

- Chemistry buffered by carbonate resulting in an alkaline soil pH



PHX



A banner for the Mars Exploration Program Analysis Group (MEPAG). The background is a dark blue, textured surface, possibly representing a Martian landscape or a satellite view of the planet's surface. On the right side, a portion of the reddish-orange planet Mars is visible, showing its characteristic color and some surface features.

Mars Exploration Program Analysis Group (MEPAG)

chartered by NASA HQ to assist in planning the scientific exploration of Mars

A large, light-colored, textured background image of the Mars surface, showing various craters and geological features. The image is slightly faded and serves as a backdrop for the main text.

**Given What We Have
Learned, Mars is an Even More
Compelling Exploration Target**

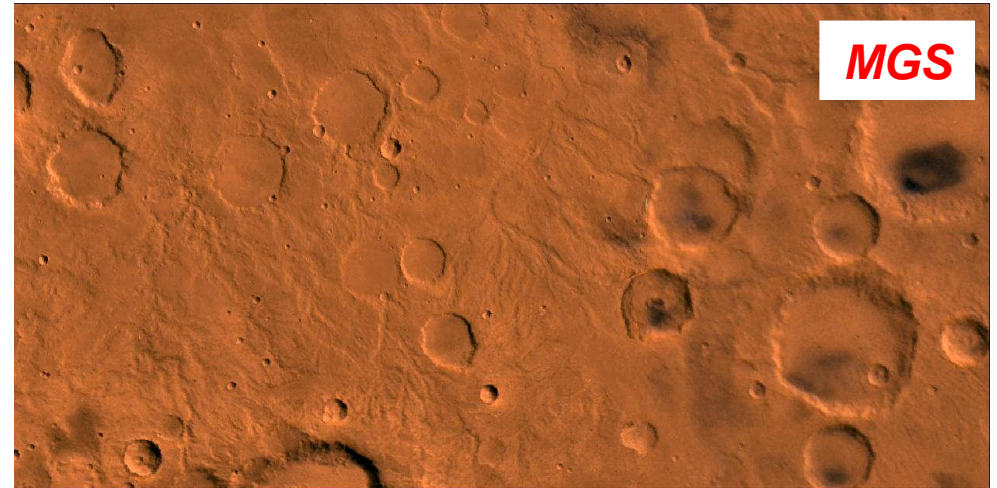
Why Mars?

- 1. Mars offers crucial information about the early evolution of the terrestrial planets, including Earth**
- 2. Mars provides a means to approach, and possibly answer, origin and evolution of life questions**
- 3. Excellent opportunity to investigate short- and long-term climate change**
- 4. Mars offers insight into the internal structure and origin of the terrestrial planets**

Why Mars?

1. Mars offers crucial information about the early evolution of the terrestrial planets, including Earth

- **Mars retains history that has been completely erased from Earth (and Venus)**
 - Earth's oldest rocks >3.5 billion years old are rare and usually altered; Mars rocks exist at 4.5 billion years (determined from dating Mars meteorites)
- ***This is the period of time when life evolved on Earth***
 - *As interpreted from chemical signatures in rock at 3.8 billion years; earliest microfossils are 3.0 billion years old.*



Ancient cratered surface of Mars (above)
and remaining Earth crust from same time period (below)



Why Mars?

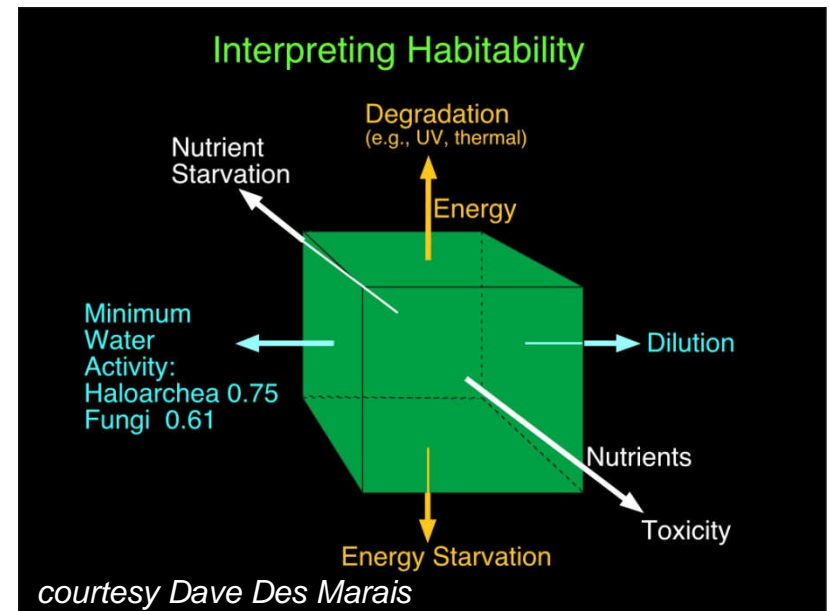
2. Mars provides a means to approach, and possibly answer, origin and evolution of life questions

❑ Ancient life—interpreted potential has increased

- Lots of ancient liquid water in diverse environments
- Past geological environments that have reasonable potential to have preserved the evidence of life, had it existed.
- Understanding variations in habitability potential is proving to be an effective search strategy
- **SUMMARY:** We have a means to prioritize candidate sites, and reason to believe that the evidence we are seeking is within reach of our exploration.

❑ Modern life—interpreted potential still exists

- Evidence of modern liquid water at surface is equivocal—probable liquid water in deep subsurface
- Methane may be a critically important clue to subsurface biosphere
- **SUMMARY:** We have not yet identified high-potential surface sites, and the deep subsurface is not yet within our reach.



Why Mars?

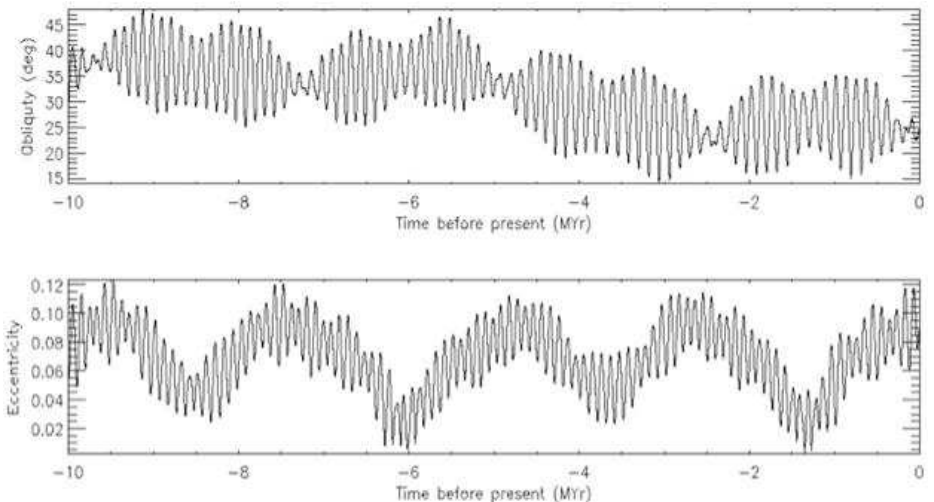
3. Excellent opportunity to investigate short- and long-term climate change

□ Preserved records of global environmental change

- Layered terrains in high- and low- latitudes indicative of cyclic changes related to orbital and axial variations
- Evidence of hydrous mineralogy changing from clays to sulfates to oxides. Mars morphology indicates water evolution over time in cooling environment.

□ Modern climate may provide clues regarding solar forcing or internal process drivers of atmospheric escape

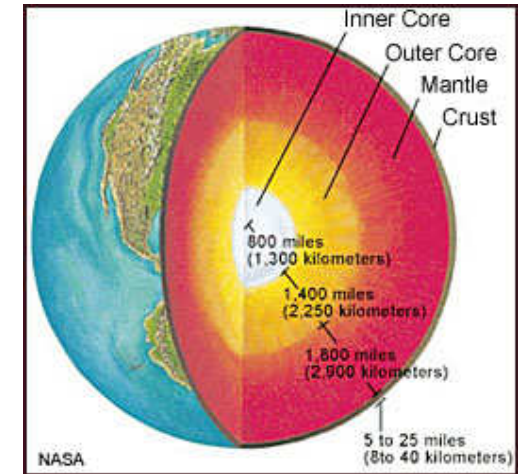
- We have observed a multi-year record of recent climate change
- The proposed MAVEN mission would establish the inventory of atmospheric trace gases to understand the internal and external processes that shaped Mars' atmosphere



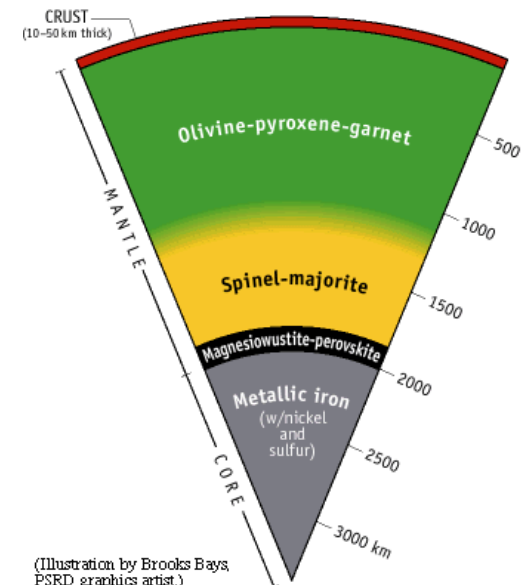
Why Mars?

4. Mars offers insight into the internal structure and origin of the terrestrial planets

- ❑ The internal structure of a planet provides clues to its origin and evolution
 - Can follow up clues from remnant magnetism discovered by MGS.
- ❑ To date, we have data for the Earth and some data for the Moon
- ❑ Mars offers an opportunity to obtain results on another terrestrial planet
 - Intermediate in size between the Earth and Moon
 - May provide clues to early differentiation that are not available from more active planets like Earth and Venus



The interior of Mars



Why Mars?

5. Strategic target for human exploration

- ❑ Closest to Earth in terms of surface environment
- ❑ Close enough that we can credibly discuss reaching it with astronauts.
- ❑ Public fascination fuels student interest in science and technology.

**Michael
Updating**

