Dynamics of braided coronal loops



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'Topological Dissipation'





Parker, E.N., ApJ, **174**, 499 (1972).

Model solar loop as braided magnetic field between parallel plates.

The model magnetic field



- Pigtail braid
- Net helicity (twist) is zero

 Could be generated by sequence of opposite-sense rotations at photosphere



Achieved in practice by adding regions of twist to uniform B Aspect ratio (1:10), low twist consistent with observed loops Conservative approach: free energy only ~3% above potential



Simulation setup

 Take field and first perform an ideal relaxation

• Then transfer to resistive MHD code: $I \times \underline{B} \approx 0$, and initialise with a uniform background plasma





t= 0.0 , J_max= 2.10





Reynolds number comparison

|**J**| at z=0

η=10-3





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Reconnection rate

Procedure, at each time:

- Integrate along an grid of field lines: $\Phi = \int E_{\parallel} ds$.
- In 2D Φ profile, seek peaks.
- Sum moduli of Φ_{max} and Φ_{min} to obtain global rec rate





Spatial coords indicate fieldline intersection with z=0 plane

Reconnection rate

- Rec rate increase a result of splitting into multiple rec regions
- (note: length of field lines = 48)

- Overall reconnected flux:
 - Total poloidal (xy) flux at t=0 : ≈ 30

at t=300: \approx 15.3

- Total reconnected flux ≈ 41.2
- Flux is multiply reconnected cf. Parnell et al. 2008

t

Rec rate comparison

- As η is decreased:
 - * No. of individual rec regions increases
 - Global rec rate varies (at most) weakly
 - * Cumulative rec'd flux increases
- Starting with 30 units poloidal flux:

η	units rec'd	units remaining
I 0 ⁻²	24.5	13.1
I 0 ⁻³	41.2	15.3
2×10-4	61.8	16.1

(grey
$$\eta = 10^{-2}$$
; dashed $\eta = 10^{-3}$; black $\eta = 2 \times 10^{-4}$)

Final state non-potential

- Initial state has net helicity zero
- Taylor relaxation would predict homogeneous B as final state
- Twist not all cancelled
- In fact final state approximates *non-linear* fff, containing flux tubes with positive and negative twist

Map of
$$\frac{J_{\parallel}}{B}$$
 at z=0

50% isosurface of

Summary

- Braided field undergoes instability, leading to 'turbulent cascade' with many reconnection events, field 'unbraids'
- Recursive reconnection of flux: poloidal flux reconnected multiple times
- For lower η:
 - * 'Cascade' enhanced in complexity and duration for lower η .
 - * Average number of reconnections increases
- Although net helicity is zero, final state does not manage to annihilate all twist (no Taylor relaxation)
- J sheets fill volume effectively (\rightarrow expect uniform heating of loop?)
- Low threshold (free energy only a few %) makes it a mechanism for 'background' heating

Pontin, D.I., Wilmot-Smith, A.L., Hornig, G., Galsgaard, K., Dynamics of Braided Coronal Loops - II. Cascade to Multiple Small Scale Events, A&A, 525, A57 (2011)

Wilmot-Smith, A.L., Pontin, D.I., Hornig, G., Dynamics of Braided Coronal Loops - I. Loss of Equilibrium, A&A, 516, A5 (2010)

Thanks for listening

References:

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- Wilmot-Smith, A.L., Pontin, D.I., Yeates, A.R. and Hornig, G. Heating of braided coronal loops, A&A, 536, A67 (2011)
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