

# Spine-Fan Reconnection

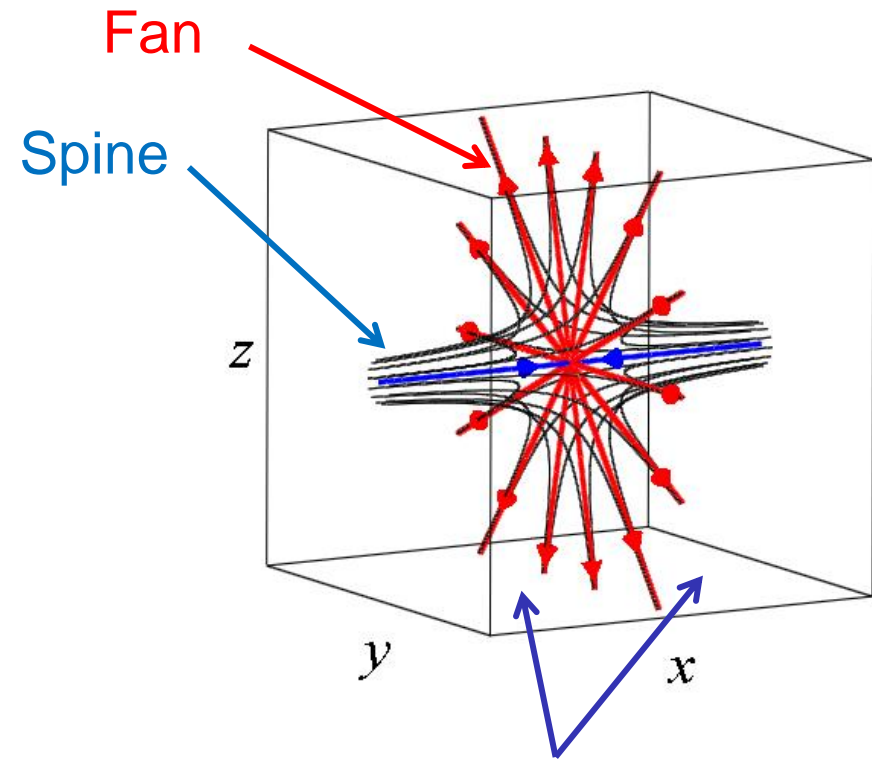
P. F. Wyper<sup>1</sup>, Rekha Jain<sup>1</sup> & D. Pontin<sup>2</sup>



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# 3D Null Points



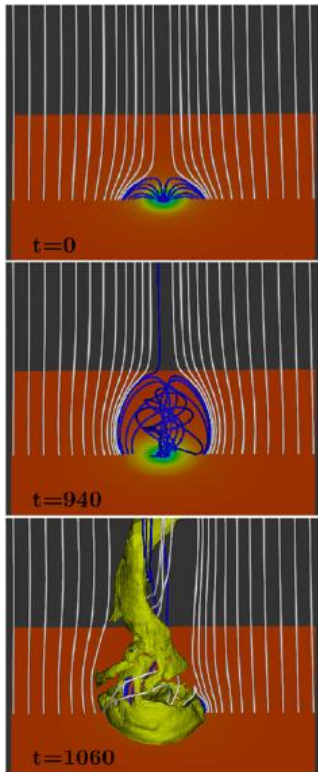
- Fields with strongly diverging field lines naturally develop current sheets<sup>1</sup>.
- The Spine & Fan of 3D nulls are one such region.

- Two distinct topological regions
- Fan plane = Separatrix surface

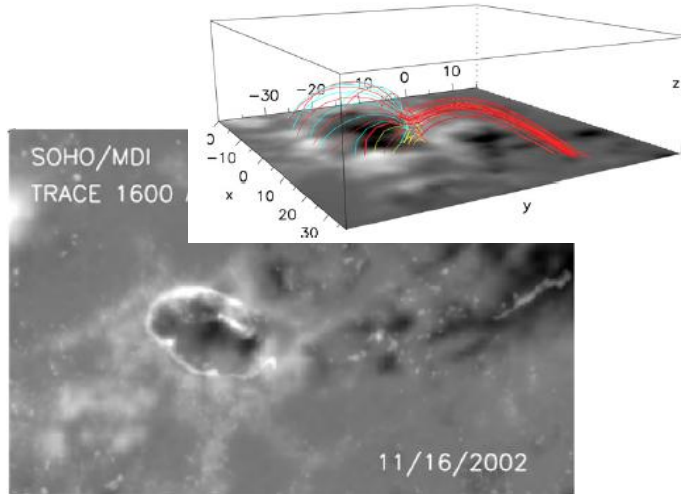
1: Priest & Démoulin (1995)

# Null Point Reconnection...

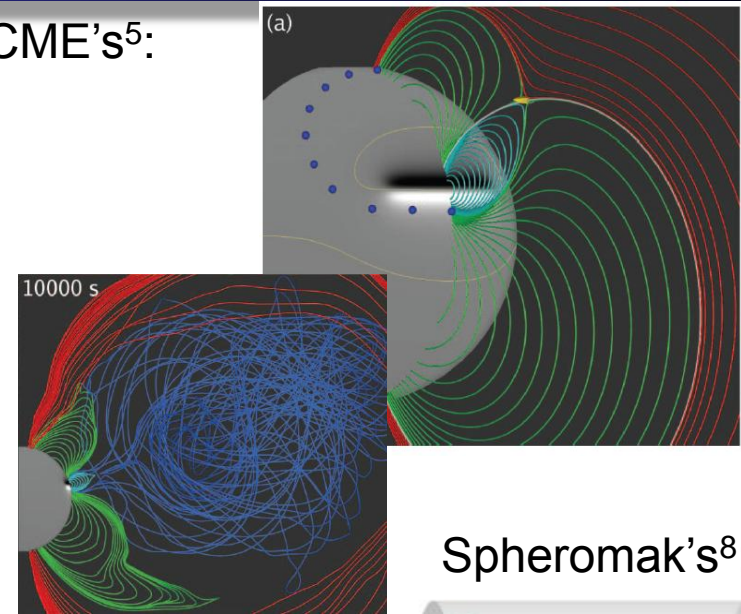
Solar Jets<sup>3</sup>



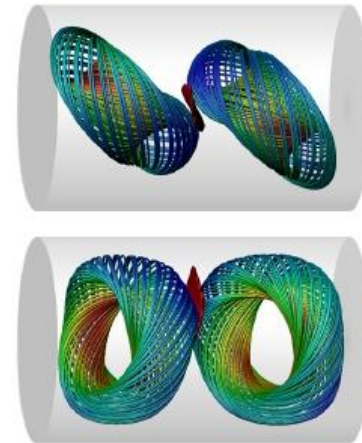
Flare brightening<sup>4</sup>



CME's<sup>5</sup>:

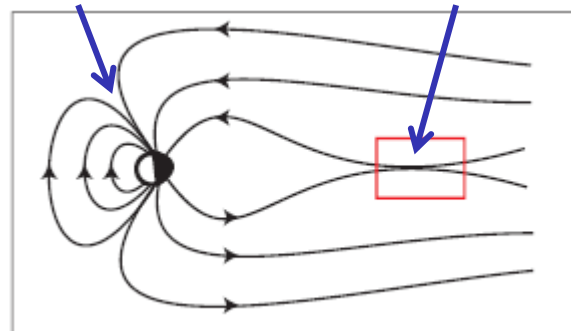


Spheromak's<sup>8</sup>:



Cusp regions<sup>2</sup>

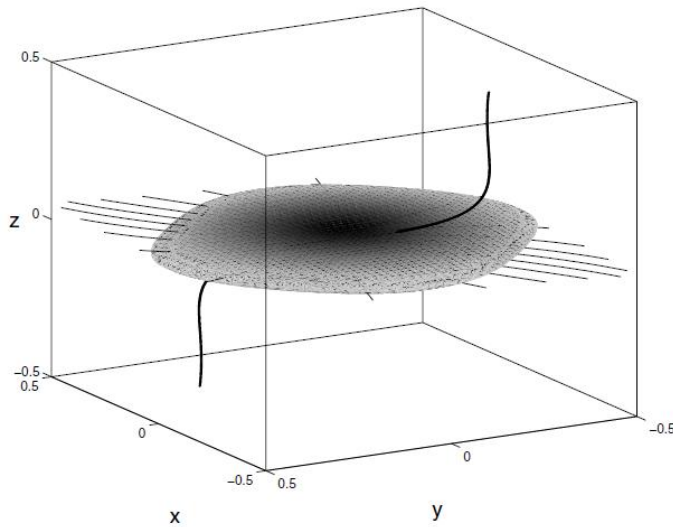
$J$  sheet in tail<sup>6,7</sup>



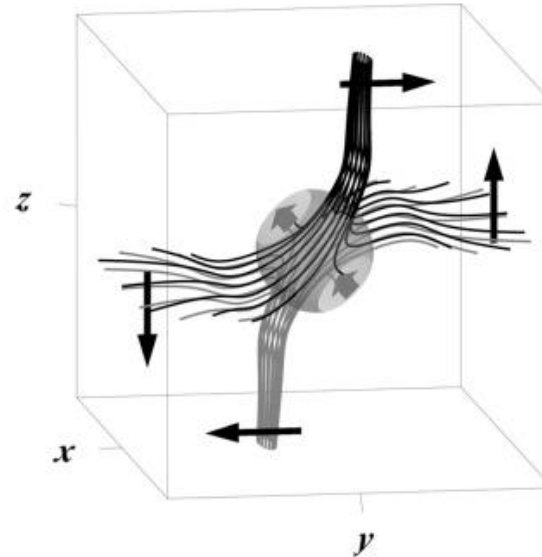
2: Dorelli et al. 07, 3: Pariat et al. 09, 4: Masson et al.,  
5: Lynch et al. 2008, 6 & 7: Xiao et al. 06,07, Gray et al. 10

# Spine-Fan

*Fan*



*Spine-Fan*



- Incompressible case:

*Exact analytical solutions<sup>8</sup>  
of reduced dimensionality*

- Compressible case:

*Transiently<sup>9</sup> and continuously<sup>10</sup>  
driven cases studied numerically*

8: Heerikhuisen & Craig (2004), 9: Priest & Pontin (2009), 10: Galsgaard & Pontin (2011)

# Questions

**All previous investigations of Spine-Fan reconnection use:**

- *Symmetric driving*
- *Driving of fixed spatial extent*

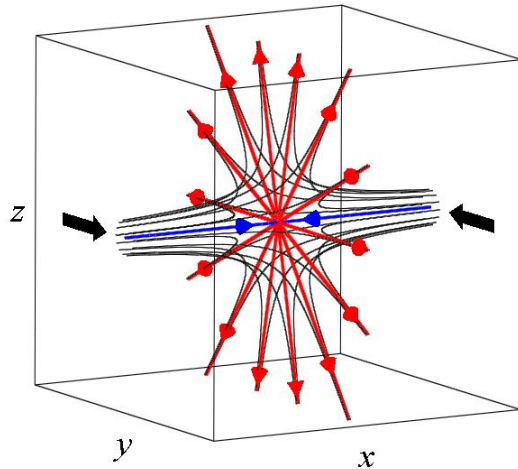
**→ Driving in the solar atmosphere is not often symmetric**

## Questions

- *How do the Spine and Fan collapse with asymmetric driving?*
- *Are the relationships seen with symmetry robust without?*
- *What is the effect on reconnection rate and current?*

# Set Up

$$\mathbf{B} = B_0(-2x, y, z)$$



- Using the Copenhagen Staggered Mesh code<sup>11,12</sup>
- Box:  $[\pm 0.5, \pm 3, \pm 3]$
- Resolution:  $128^3$
- Stretched grid:  
 $\delta x \sim 0.005, \delta y = \delta z \sim 0.025$
- *Line tied* driving boundaries
- *Closed* side boundaries with thin damping region.

• Ideal Gas:  $\gamma = 5/3$

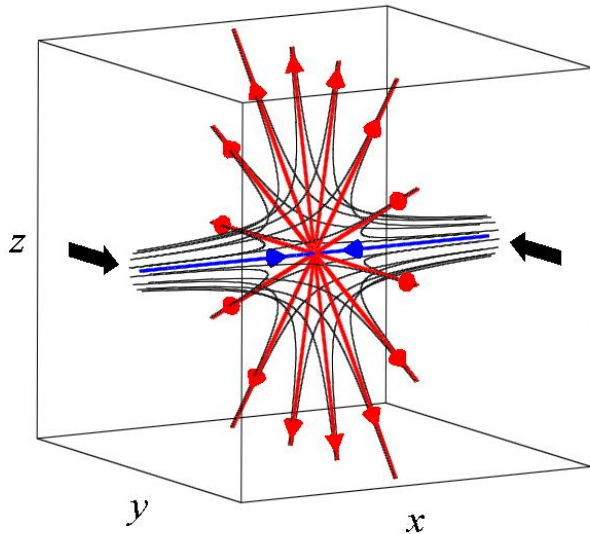
• Plasma  $\beta = \frac{10\mu_0\beta^*}{3B^2}$

•  $\beta^* = 0.05, B_0 = 1$

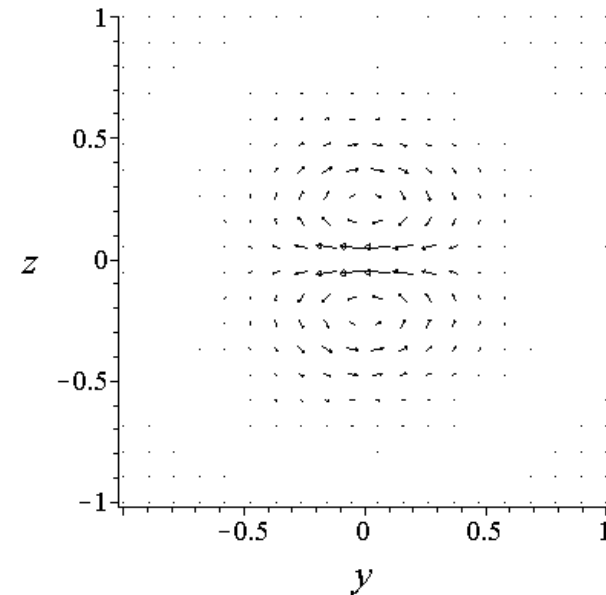
11: Galsgaard & Nordlund (1997), 12: <http://www.astro.ku.dk/~kg>

# Driver

$$\mathbf{B} = B_0(-2x, y, z)$$

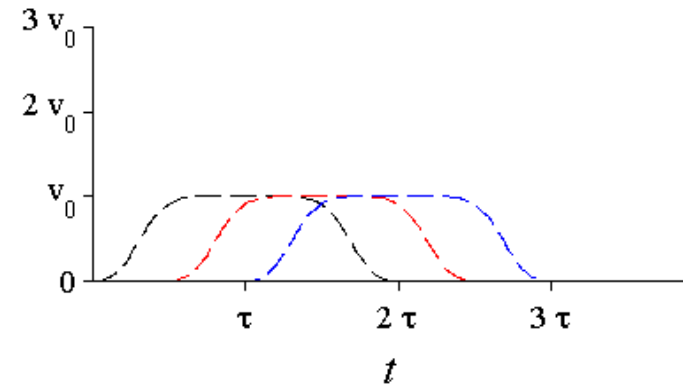


## Driving patch



Time dependence:

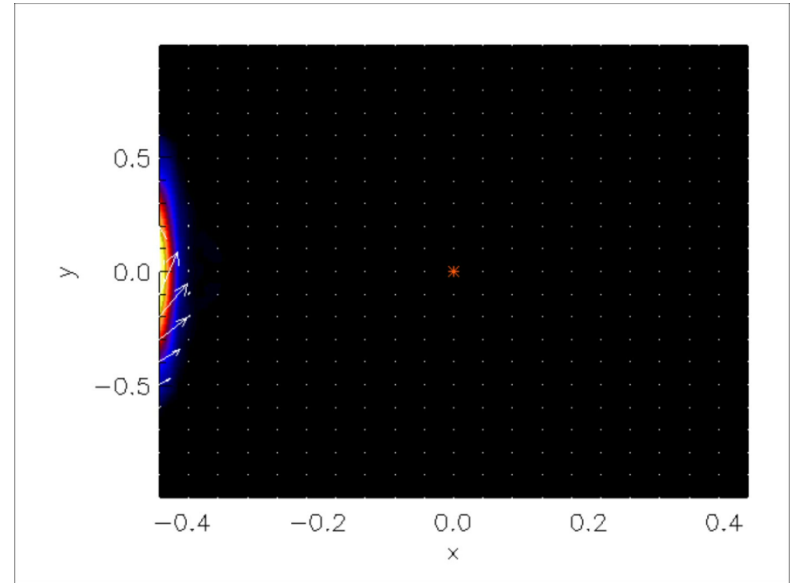
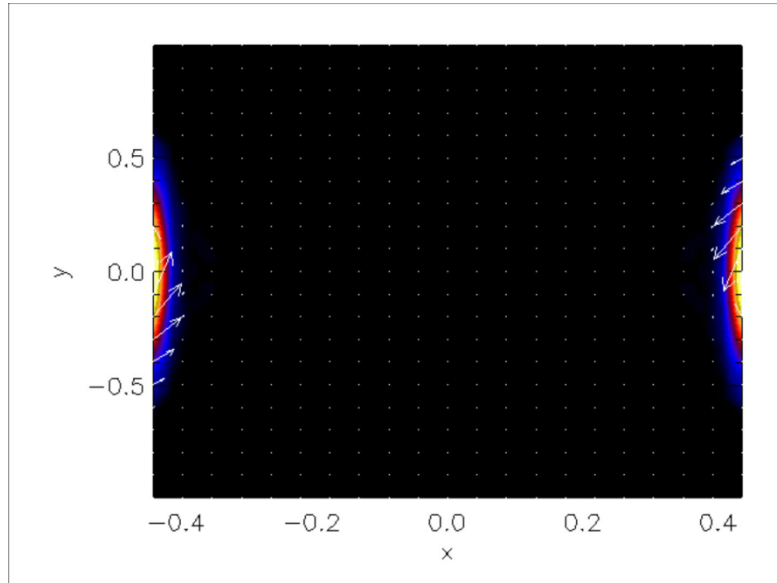
$$V_0(t) = v_0 \tanh(t/0.1)^2 \begin{cases} e^{-0.2(t-\tau)^2}, & x = -0.5 \\ e^{-0.2(t-\tau-t_{lag})^2}, & x = 0.5 \end{cases}$$



13: Pontin *et al.* (2007)

# One vs Two Pulses

- $z = 0$  plane with  $\eta = 5 \times 10^{-4}$ ,  $\tau = 1.8$ ,  $v_0 = 0.04$



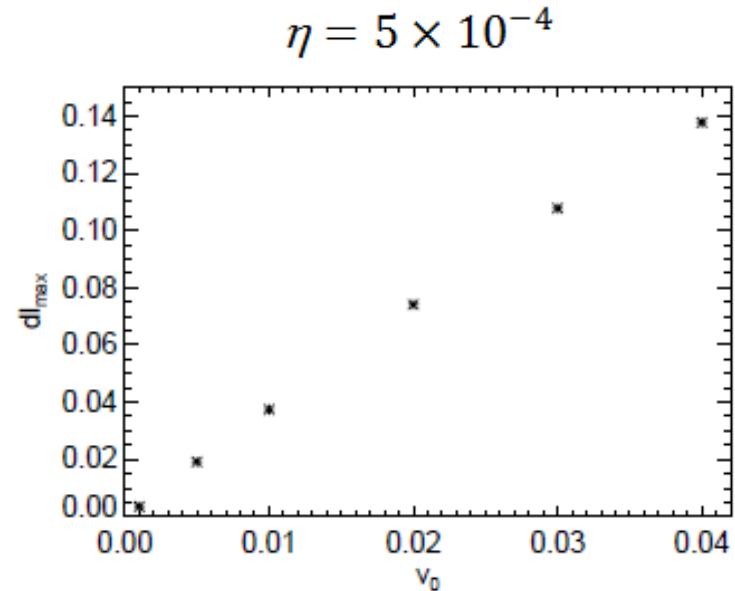
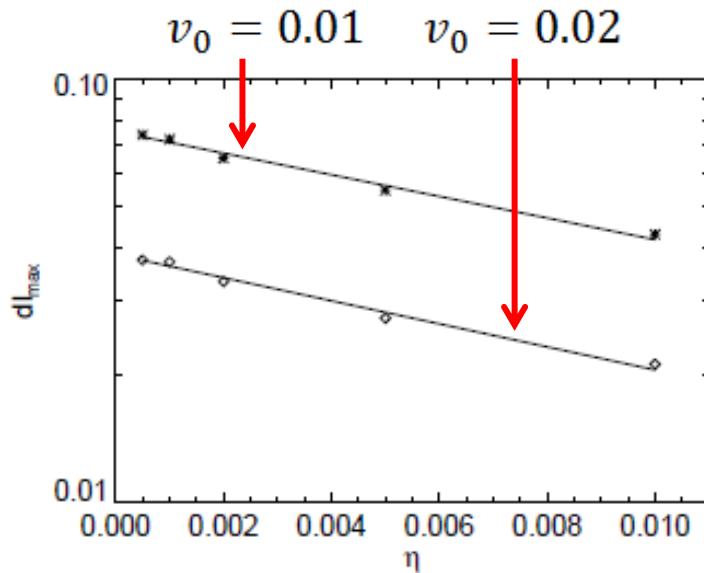
- Null position maintained
- Spine & Fan crossing flows
- Null position shifts
- Spine moves with null



# Null Displacement

Null displacement relationships:

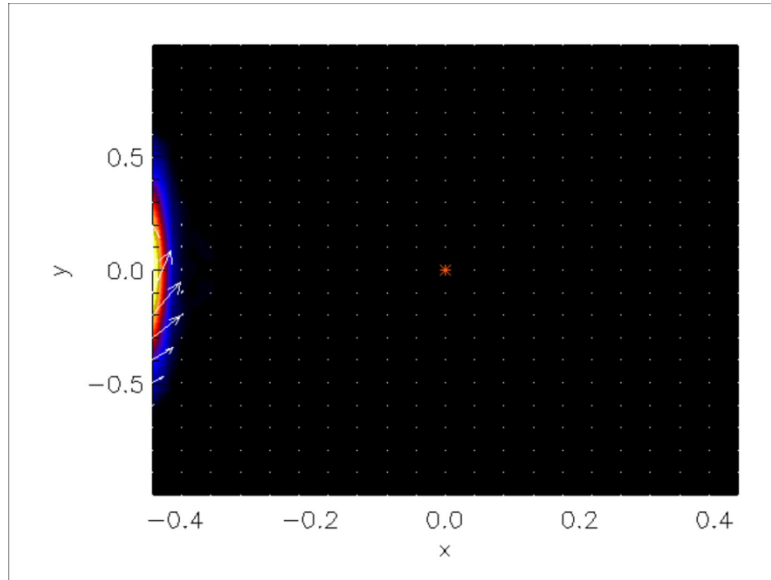
$$dl = \sqrt{x_{null}^2 + y_{null}^2}$$



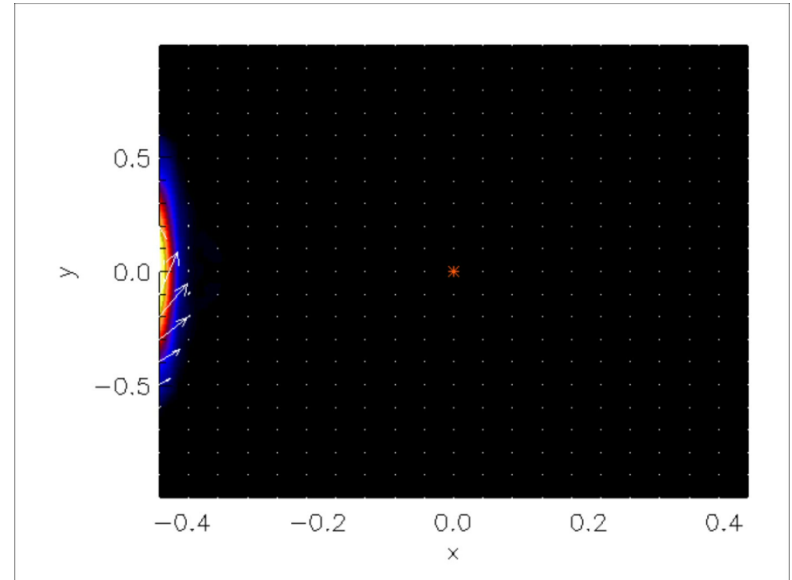
- $dl_{max}$  increases linearly with  $v_0$  and decreases exponentially with  $\eta$ .
- Smaller  $\eta$  = more collapse (stronger  $\mathbf{J} \times \mathbf{B}$ )

# Pulses separated by a lag period

- $z = 0$  plane with  $\eta = 5 \times 10^{-4}$ ,  $\tau = 1.8$ ,  $v_0 = 0.04$



$$t_{lag} = 0.9$$

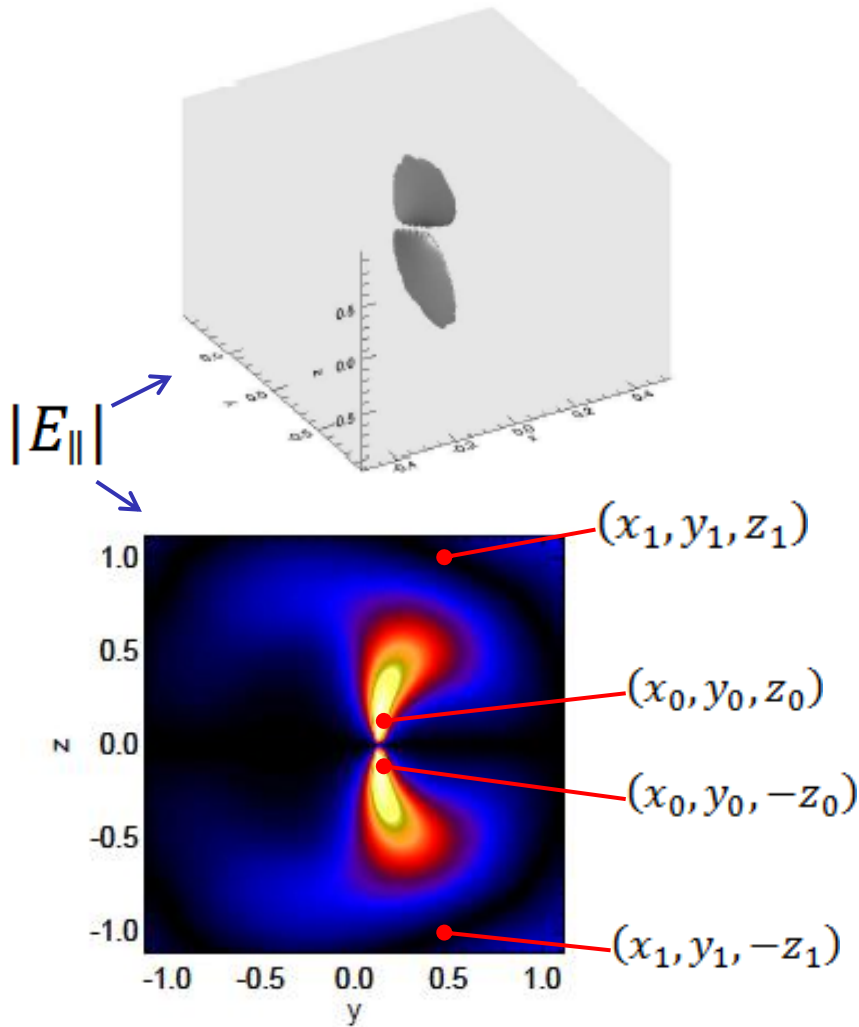


$$t_{lag} = 1.8$$

- Greatest Spine-Fan collapse when pulses best overlap.
  - Null shifts back and forth with the pulses

# Reconnection Rate

- Field line of max  $E_{\parallel}$  must be found iteratively:



$$F = \max \left\{ \int_{(x_0, y_0, z_0)}^{(x_1, y_1, z_1)} E_{\parallel} dl + \int_{(x_1, y_1, -z_1)}^{(x_0, y_0, -z_0)} E_{\parallel} dl \right\}$$

+null contribution

$$= \max \left\{ \int_{(x_1, y_1, -z_1)}^{(x_1, y_1, z_1)} E_{\parallel} dl \right\}$$

- The values of each point change at every step.
- $z_0 = 0.05$  throughout.

14: Pontin *et al.* (2005)

# Variation in Time

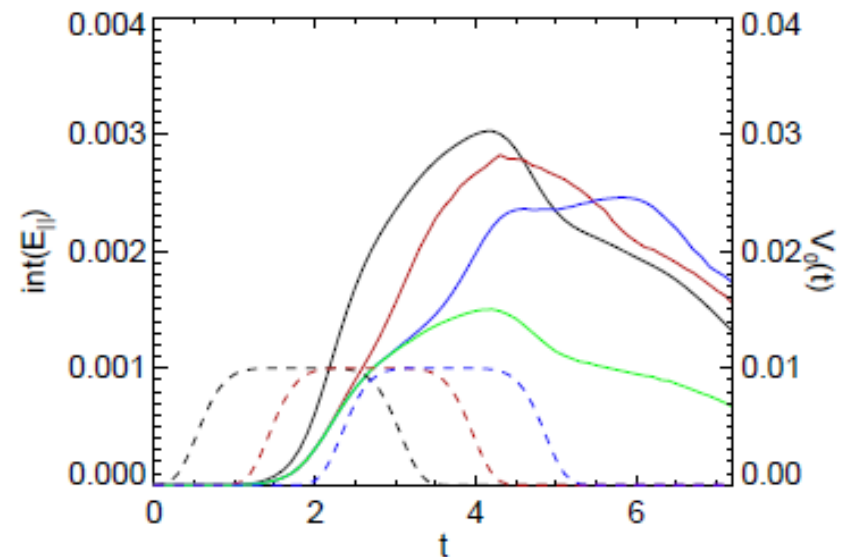
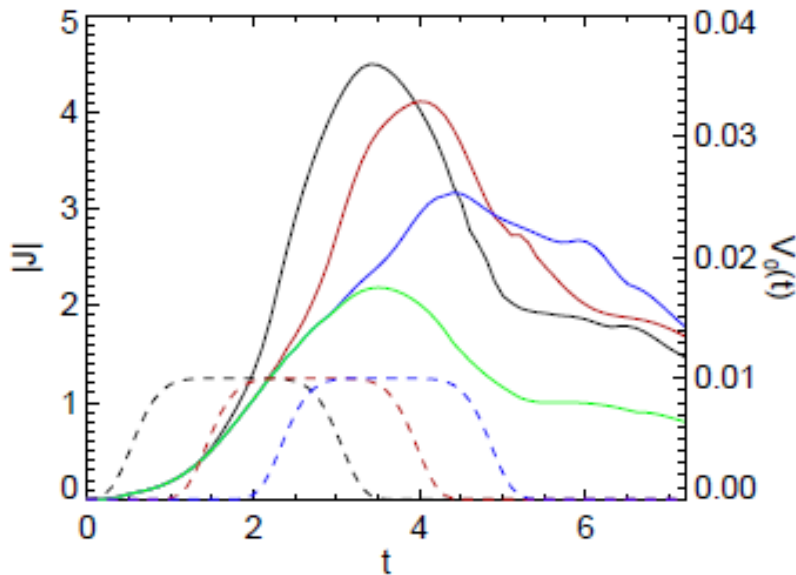
Green: *One pulse*

Black:  $t_{lag} = 0$

$$v_0 = 0.01$$

Red:  $t_{lag} = 0.9$

Blue:  $t_{lag} = 1.8$



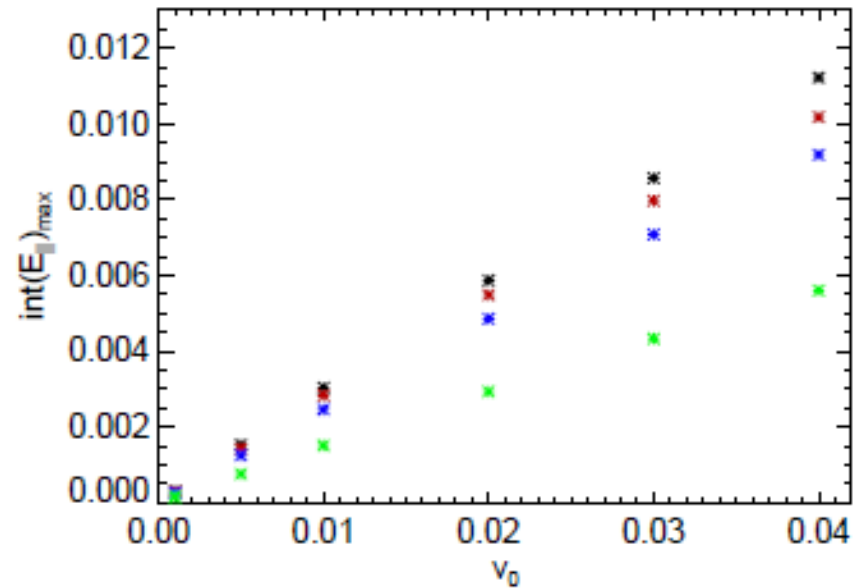
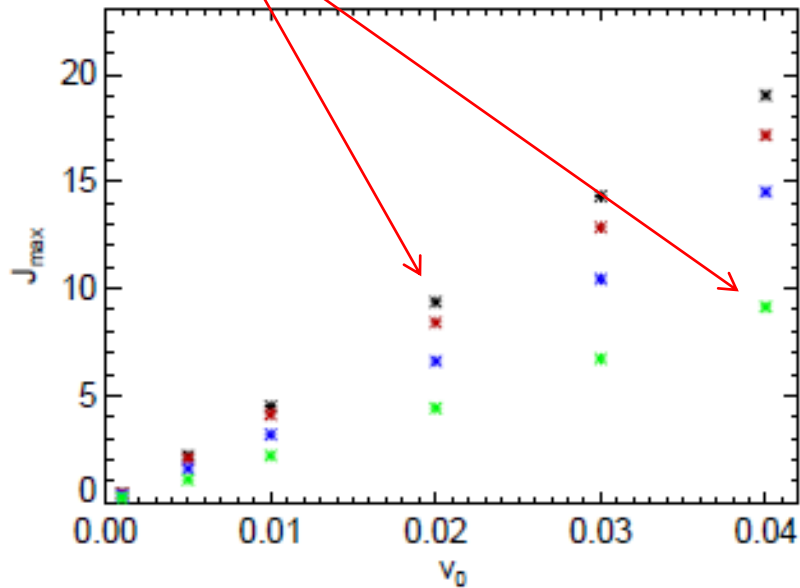
- Peaks occurs when the latent shear from the 1<sup>st</sup> pulse best overlaps the 2<sup>nd</sup>

- Peaks later than current
- Difference in overlaps

# Qualitative Trends

*Gradient is half*

*Linear trends remain*

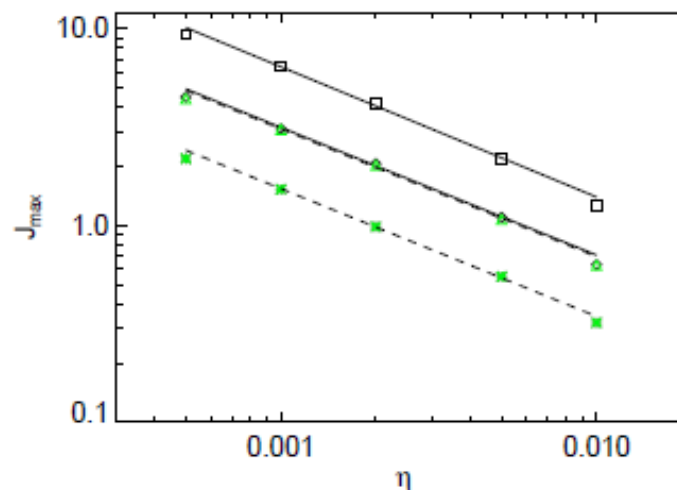
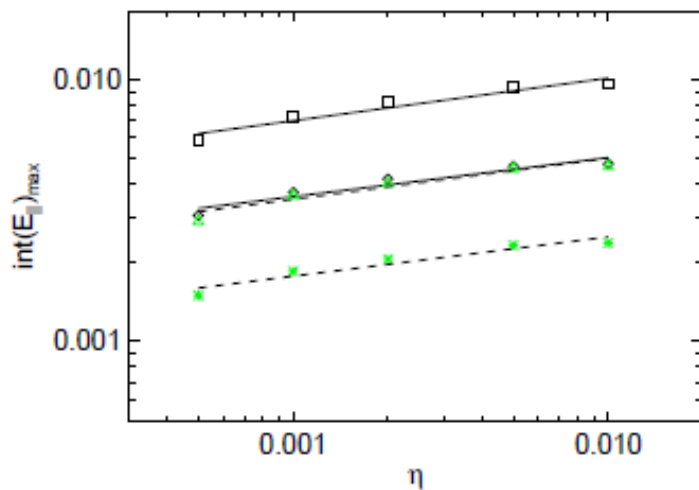


# Resistive Scalings

Does the null shifting affect scaling relations?<sup>15,16,17</sup>

Green: *One pulse* Black: Two symmetric pulses

Dashed:  $v_0 = 0.01$  Solid:  $v_0 = 0.02$



- Scaling relations not changed by null movement

$$\left[ \int E_{\parallel} dl \right]_{\text{max}} \propto \eta^{0.15}$$

$$|J|_{\text{max}} \propto \eta^{-0.65}$$

15: Lukin & Linton (2011), 16: Heerikhuisen & Craig (2004), 17: Galsgaard & Pontin (2011)

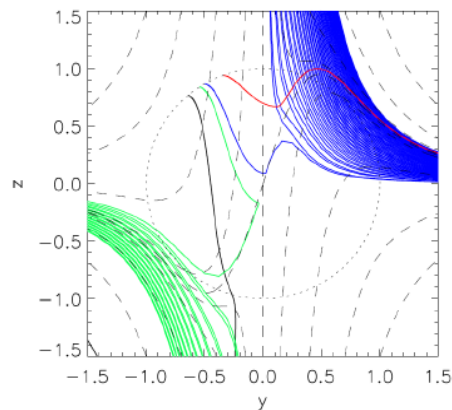
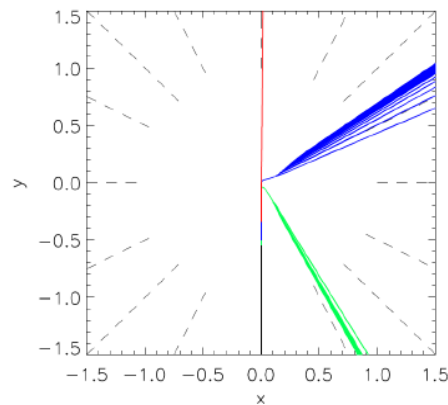
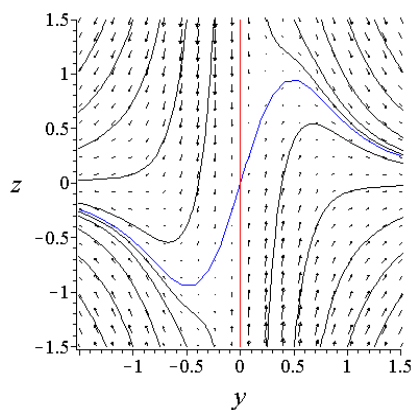
# Conclusions

## Questions/**Answers**

- **How does Spine-Fan collapse with asymmetric driving?**  
*Asymmetrically with null movement depending on  $\eta$  and  $v_0$ .*
- **Are the relationships seen with symmetry robust without?**  
*Linear trends in  $J_{max}$  and  $\text{int}(E_{\parallel})_{max}$  with  $v_0$  robust.  $\eta$  scaling independent of null movement.*
- **What is the effect on reconnection rate and current?**  
*Heavily dependent on the timing and size of the driving patches.*

# Future Work

- *Particle acceleration?*
- *Non-potential initial fields?*
- *Sheet fragmentation?*
- *Small scale physics?*

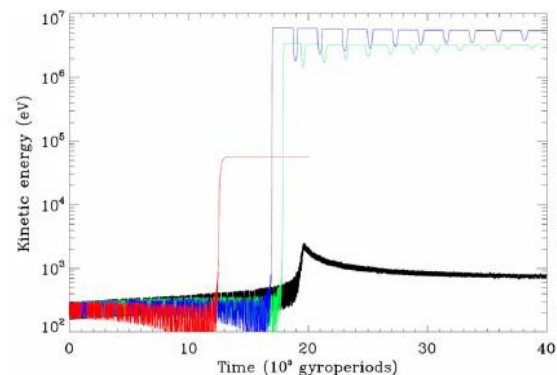


$$\vec{B} = \frac{B_0}{L} (x, y, -2z + j_0 y \exp(-4R^2))$$

$$\Phi = -\int \eta \vec{J} \cdot \vec{B} ds$$

$$\frac{d\vec{x}}{dt} = \frac{\vec{p}}{\gamma m}$$

$$\frac{d\vec{p}}{dt} = q \left( \vec{E} + \frac{1}{\gamma m} [\vec{p} \times \vec{B}] \right)$$



18: Dalla & Browning (2005)