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An improved $k - \omega$ model for 3D hypersonic flows

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Context & Equations Context

2 Theoretical and numerical solvings

3 Results

- Ramp at mach 5 (Delery 1990)
- Dutton and Herrin experiment

4 Conclusions & Prospects





Aim

- Our aim is to capture fluxes (pressure and thermal fluxes) in the flow over a blunt hypersonic body.
- Turbulence has a great impact on the values of these quantities so we need a robust and accurate model to do the job.
- Here we present a first-order turbulent model based on Wilcox $k \omega$ model.
- We are only interested in solving stationnary states.

Cea Incompressible $k - \omega$ turbulent equations

Equations for U (velocity), k (turbulent kinetic energy) and ω (characteristic frequency of the turbulence)

$$\begin{split} \partial_t U + U \cdot \nabla U + \nabla \left(P + \frac{2}{3} k \right) &= \nabla \cdot \left[\left(\nu + \nu_t \right) \left(\nabla U + \nabla U^T \right) \right], \\ \nabla \cdot U &= 0, \\ \partial_t k + U \cdot \nabla k &= 2\nu_t \left(S : S \right) - \beta^* k \omega + \nabla \cdot \left[\left(\nu + \sigma^* \nu_t \right) \nabla k \right], \\ \partial_t \omega + U \cdot \nabla \omega &= 2\alpha \left(S : S \right) - \beta \omega^2 + \nabla \cdot \left[\left(\nu + \sigma \nu_t \right) \nabla \omega \right], \\ \nu_t &= \frac{k}{\omega}, \alpha = \frac{13}{25}, \beta = \beta_0 f_\beta, \ \beta^* &= \beta_0^* f_\beta, \\ \sigma &= \sigma^* = 0.5, \beta_0 = \frac{9}{125}, \ \beta_0^* &= \frac{9}{100}, \\ f_\beta &= \frac{1 + 70\chi_\omega}{1 + 80\chi_\omega} \text{ où } \chi_\omega = \left| \frac{(\Omega \otimes \Omega) : S}{(\beta_0 \omega)^3} \right|, \\ f_\beta^* &= \begin{cases} 1 & , & \chi_k \leq 0 \\ \frac{1 + 680\chi_k^2}{1 + 400\chi_k}, & \chi_k \geq 0 \end{cases}, \ \chi_k &= \frac{1}{\omega^3} \nabla k \nabla \omega. \\ \text{CEACESTA | Multimat 2013, San Francisco | PAGE 4/32 \end{cases} \end{split}$$

From incompressible flows towards compressible flows

Compressible equations

- Compressible equations are obtained by analogy with incompressible equations. Mean quantities are defined by pondering by the mass (Favre's averaging).
- We ensure in the code that Reynolds Tensor T

 $R = (\nu + \nu_t) \left(\nabla U + \nabla U^T \right) - \frac{2}{3} k I$ have all his diagonal coefficients

negative and all non diagonal satisfy $|R_{ij}| \le \sqrt{R_{ii}R_{jj}}$ because of its physical meaning ("realisibility condition").

Cea Final system of equations (1)

$$\begin{split} \partial_t(\rho) + \nabla \cdot (\rho U) &= 0, \\ \partial_t(\rho U) + \nabla (\rho U \otimes U) + \nabla \left(P + \frac{2}{3}\rho k\right) \\ &= \nabla \cdot \left[\rho \left(\nu + \nu_t\right) \left(\nabla U + \nabla U^T - \frac{2}{3}(\nabla \cdot U)\mathbf{I}\right)\right], \\ \partial_t(\rho E) + \nabla \left(\left(\rho E + P + \frac{2}{3}\rho k\right)U\right) - \nabla \cdot \left(\left(\lambda + \frac{\rho \nu_t C_p}{Pr_t}\right)\nabla T\right) \\ &= \nabla \cdot \left[\rho \left(\nu + \nu_t\right) \left(\nabla U + \nabla U^T - \frac{2}{3}(\nabla \cdot U)\mathbf{I}\right)U\right] + \nabla \cdot \left[\rho \left(\nu + \sigma^* \nu_t\right)\nabla k\right], \\ \partial_t(\rho k) + \nabla \cdot (\rho k U) &= -f_c \rho R : \nabla U - \beta^*_{new} \rho k \omega + \nabla \cdot \left[\rho \left(\nu + \sigma^* \nu_t\right)\nabla k\right], \\ \partial_t(\rho \omega) + \nabla \cdot (\rho \omega U) &= -\alpha f_c \rho R : \nabla U - \beta_{new} \rho \omega^2 + \nabla \cdot \left[\rho \left(\nu + \sigma \nu_t\right)\nabla \omega\right], \\ \nu_t &= \alpha^* \frac{k}{\omega}, \nu = \frac{\mu}{\rho}, \\ S &= \frac{\nabla U + \nabla U^T}{2}, \Omega = \frac{\nabla U - \nabla U^T}{2}, R = \nu_t \left(2S - \frac{2}{3}(\nabla \cdot U)\mathbf{I}\right) - \frac{2}{3}k\mathbf{I}, \end{split}$$

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Cea Final system of equations (2)

$$\begin{split} &f_{c} = 1.7, Re_{t} = \frac{k}{\nu\omega}, \sigma = \sigma^{*} = 0.5, \beta_{0} = \frac{9}{125}, \beta_{0}^{*} = \frac{9}{100}, \alpha_{0}^{*} = \frac{1}{3}\beta_{0}, \alpha_{0} = \frac{1}{9}\\ &\alpha^{*} = \frac{\alpha_{0}^{*} + Re_{t}/R_{k}}{1 + Re_{t}/R_{k}}, \alpha = \frac{13}{25}\frac{\alpha_{0} + (Re_{t}/R_{\omega})}{1 + (Re_{t}/R_{\omega})}(\alpha^{*})^{-1},\\ &\beta = \beta_{0}\frac{4/15 + (Re_{t}/R_{\beta})^{4}}{1 + (Re_{t}/R_{\beta})^{4}}f_{\beta}, \beta^{*} = \beta_{0}^{*}f_{\beta},\\ &\beta^{*}_{new} = \beta^{*} - 1.5\max(Ma_{t}^{2} - Ma_{t_{0}}^{2}, 0)\beta \qquad \text{où} \qquad Ma_{t_{0}} = 0.5\\ &\beta_{new} = \beta(1 + 1.5\max(Ma_{t}^{2} - Ma_{t_{0}}^{2}, 0))\\ &R_{\beta} = 8, R_{k} = 6, R_{\omega} = 6\\ &f_{\beta} = \frac{1 + 70\chi\omega}{1 + 80\chi\omega} \qquad \text{où} \qquad \chi_{\omega} = \left|\frac{(\Omega \otimes \Omega) : S}{(\beta_{0}\omega)^{3}}\right|,\\ &f_{\beta}^{*} = \begin{cases} 1 & , & \chi_{k} \leq 0\\ \frac{1 + 680\chi_{k}^{2}}{1 + 400\chi_{k}}, & & \chi_{k} \geq 0 \end{cases}, \quad \chi_{k} = \frac{1}{\omega^{3}}\nabla k\nabla\omega,\\ ⪻_{t} = 0.7. \end{cases}$$



Boundary layer equations

$$\begin{aligned} \partial_t(\rho k) + \nabla \cdot (\rho k U) &= \rho R : \nabla U - \beta^* \rho k \omega + \nabla \cdot [\rho \left(\nu + \sigma^* \nu_t\right) \nabla k], \\ \partial_t(\rho \omega) + \nabla \cdot (\rho \omega U) &= \alpha \rho R : \nabla U - \beta \rho \omega^2 + \nabla \cdot [\rho \left(\nu + \sigma \nu_t\right) \nabla \omega]. \end{aligned}$$

If one looks for $k \sim C_k y^n$ and $\omega \sim \frac{C_\omega}{y^n}$ so that dissipation $\epsilon = k\omega$ tends to a finite value when the wall distance *y* tends to zero, one gets:

$$0 = -\beta^* C_{\omega} C_k + \nu (n(n-1)) C_k y^{n-2},$$
 (1)

$$0 = -\beta C_{\omega}^2 y^{-2n} + \nu (n(n+1)) C_{\omega} y^{-n-2}.$$
 (2)

Equation (1) gives n = 2 and $\beta^* C_\omega = 2\nu$. Zquation (2) is satisfied if and only if -2n = n - 2 which also means n = 2 and $\beta C_\omega = 6\nu$.





Context & Equations Context

2 Theoretical and numerical solvings

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Some mathematical properties of a simplified incompressible model(1)

$$\begin{aligned} \partial_t U + U \cdot \nabla U + \nabla \left(P + \frac{2}{3} k \right) &= \nabla \cdot \left[\left(\nu + \nu_t \right) \left(\nabla U + \nabla U^T \right) \right], \\ \nabla \cdot U &= 0, \\ \partial_t k + U \cdot \nabla k &= 2\nu_t \left(S : S \right) - \beta^* k \omega + \nabla \cdot \left[\left(\nu + \sigma^* \nu_t \right) \nabla k \right], \\ \partial_t \omega + U \cdot \nabla \omega &= 2\alpha \left(S : S \right) - \beta \omega^2 + \nabla \cdot \left[\left(\nu + \sigma \nu_t \right) \nabla \omega \right], \\ \nu_t &= \frac{k}{\omega}, \alpha = \frac{13}{25}, \sigma = \sigma^* = 0.5, \beta = \frac{9}{125}, \beta^* = \frac{9}{100}. \end{aligned}$$

If one looks for regular local in time 3D solutions to the $k - \omega$ model (in Sobolev spaces) on bounded regular domains we have the following properties:

Framework of the work

- we suppose that U is in the Sobolev space H_0^s for small times with s > 4 + 3/2.
- we suppose that k is in the Sobolev space H_0^s for small times with s > 4 + 3/2. item we suppose that $\omega \omega_0$ is in the Sobolev space H_0^s for small times with s > 4 + 3/2 where ω_0 is a constant (the boundary condition on the domain).

Some properties of a simplified incompressible model(2)

Then we have

A priori estimates

There exists some integer *n* depending on *s* such that

$$\frac{d}{dt}\left(||U||^{2}_{H^{s}}+||k||^{2}_{H^{s}}+||\omega-\omega_{0}||^{2}_{H^{s}}\right) \leq \left(||U||^{2}_{H^{s}}+||k||^{2}_{H^{s}}+||\omega-\omega_{0}||^{2}_{H^{s}}\right)^{n(s)}$$

so that if one finds ways to construct solutions through an iterative process there will exist local in time smooth solutions in 3D. (Mathiaud 2008 :"Local smooth solutions of the incompressible $k - \epsilon$ model and the low turbulent diffusion limit)



The global system of compressible equations can be under the following form:

$$\partial_t V + \partial_x F(V) + \partial_{xx}^2 G(V) = S(V)$$
(3)

where *V* is the vector $\begin{pmatrix} \rho U \\ \rho E \\ \rho k \\ \rho \omega \end{pmatrix}$, *F*(*V*) represents the flux associated to *V*, G(V) the diffusive part of the system and S(V) the source term.

The time marching is solved through:

$$\frac{V^{n+1} - V^n}{dt} + \partial_x F(V^{n+1}) + \partial_{xx}^2 G(V^{n+1}) = S(V^{n+1})$$
(4)

We use a finite volume scheme for Navier-Stokes equations with a Roe-type solver.

We note

$$\Delta(\rho k) = \rho^{n+1} k^{n+1} - \rho^n k^n, \qquad (5)$$

$$\Delta(\rho\omega) = \rho^{n+1}\omega^{n+1} - \rho^n\omega^n, \qquad (6)$$

$$\Delta \rho = \rho^{n+1} - \rho^n, \tag{7}$$

$$\Delta(\rho U) = \rho^{n+1} U^{n+1} - \rho^n U^n.$$
(8)

Implicitation of k and ω equations $\mathcal{O}\mathcal{O}$

If one linearizes turbulent kinetic equation around the state at time *n* one gets:

$$\begin{aligned} \frac{\Delta\rho k}{dt} + \nabla\cdot\left(U^{n}\rho^{n}k^{n}\right) + \nabla\cdot\left(U^{n}\Delta\rho k\right) + \nabla\cdot\left(k^{n}\Delta\rho U\right) + \nabla\cdot\left(-\frac{U^{n}k^{n}}{\left(\rho^{n}\right)^{2}}\Delta\rho\right) \\ = & \left(-f_{c}\rho R:\nabla U\right)^{n} - \left(\beta_{new}^{*}\right)^{n}\rho^{n}k^{n}\omega^{n} - \left(\beta_{new}^{*}\right)^{n}\omega^{n}\Delta\rho k - \left(\beta_{new}^{*}\right)^{n}k^{n}\Delta\rho\omega \\ + & \nabla\cdot\left[\left(\rho\left(\nu+\sigma^{*}\nu_{t}\right)^{n}\nabla k^{n}\right] + \nabla\cdot\left[\left(\nu+\sigma^{*}\nu_{t}\right)^{n}\nabla\left(\Delta\rho k\right)\right] \\ - & \nabla\cdot\left[\left(\left(\nu+\sigma^{*}\nu_{t}\right)^{n}k^{n}\right)\nabla\left(\Delta\rho\right)\right] \end{aligned}$$

Then turning to ω equation:

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$$\begin{aligned} \frac{\Delta\rho\omega}{dt} + \nabla\cdot\left(U^{n}\rho^{n}\omega^{n}\right) + \nabla\cdot\left(U^{n}\Delta\rho\omega\right) + \nabla\cdot\left(\omega^{n}\Delta\rho U\right) + \nabla\cdot\left(-\frac{U^{n}\omega^{n}}{\left(\rho^{n}\right)^{2}}\Delta\rho\right) \\ = & \left(-\alpha f_{c}\rho R:\nabla U\right)^{n} - \left(\beta_{new}^{*}\right)^{n}\rho^{n}\omega^{n}\omega^{n} - \left(\beta_{new}^{*}\right)^{n}\omega^{n}\Delta\omega - \left(\beta_{new}^{*}\right)^{n}\omega^{n}\Delta\omega \\ + & \nabla\cdot\left[\left(\rho\left(\nu+\sigma^{*}\nu_{t}\right)\right)^{n}\nabla\omega^{n}\right] + \nabla\cdot\left[\left(\nu+\sigma^{*}\nu_{t}\right)^{n}\nabla\left(\Delta\rho\omega\right)\right] \\ - & \nabla\cdot\left[\left(\left(\nu+\sigma^{*}\nu_{t}\right)^{n}\omega^{n}\right)\nabla\left(\Delta\rho\right)\right] \end{aligned}$$

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Looking for stationnary states

$$\frac{\Delta\rho k}{dt} + \nabla \cdot (\boldsymbol{U}^{n}\rho^{n}k^{n}) + \nabla \cdot (\boldsymbol{U}^{n}\Delta\rho k) + \nabla \cdot (k^{n}\Delta\rho U) + \nabla \cdot \left(-\frac{\boldsymbol{U}^{n}k^{n}}{(\rho^{n})^{2}}\Delta\rho\right)$$
$$= (-f_{c}\rho\boldsymbol{R}:\nabla\boldsymbol{U})^{n} - (\beta_{new}^{*})^{n}\rho^{n}k^{n}\omega^{n} - (\beta_{new}^{*})^{n}\omega^{n}\Delta\rho k - (\beta_{new}^{*})^{n}k^{n}\Delta\rho\omega$$
$$+ \nabla \cdot \left[(\rho(\nu + \sigma^{*}\nu_{t}))^{n}\nabla k^{n}\right] + \nabla \cdot \left[(\nu + \sigma^{*}\nu_{t})^{n}\nabla(\Delta\rho k)\right]$$

$$\frac{\Delta\rho\omega}{dt} + \nabla \cdot (\boldsymbol{U}^{n}\rho^{n}\omega^{n}) + \nabla \cdot (\boldsymbol{U}^{n}\Delta\rho\omega) + \nabla \cdot (\omega^{n}\Delta\rho U) + \nabla \cdot \left(-\frac{\boldsymbol{U}^{n}\omega^{n}}{\left(\rho^{n}\right)^{2}}\Delta\rho\right)$$

$$= (-\alpha f_c \rho \mathbf{R} : \nabla U)^n - (\beta_{new}^*)^n \rho^n \omega^n \omega^n - (\beta_{new}^*)^n \Delta \omega - (\beta_{new}^*)^n \omega^n \Delta \omega + \nabla \cdot [(\rho (\nu + \sigma^* \nu_t))^n \nabla \omega^n] + \nabla \cdot [(\nu + \sigma^* \nu_t)^n \nabla (\Delta \rho \omega)]$$





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Cea 2D ramp at mach 5 (Delery 1990)



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Figure: Pressure coefficient & Stanton Number

$$C_{p} = \frac{P - P_{\infty}}{\frac{1}{2}\rho_{\infty}V_{\infty}^{2}}, \qquad St = \frac{\lambda \nabla T \cdot \vec{r}}{\rho_{\infty}V_{\infty}C\left(T_{\infty}(1 + \frac{\gamma - 1}{2}M_{\infty}^{2}) - T_{p}\right)}$$

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Figure: Pressure coefficient & Stanton Number

3D results are quite in agreement with 2D results

Dutton and Herrin experiment(1994, AIAA Journal Vol. 32, No. 1, January)



Mach = 2.46, Pressure = 32078.5, Density =0.84302, Velocity= 567.8

Cea Mach Number and Turbulent intensity



Figure: Recirculation point measured at 0.084m by Dutton and assessed at 0.1m by the code



Ce2 Pressure results for different models



Figure: C_p on the wall

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Cea 2D (5.10^4 cells) .vs. 3D (10^7 cells)



Figure: C_p on the wall

Figure: Reattachment point



Putting rough boundary conditions



Figure: Reattachment point

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Flaws in the mesh creating orthoradial velocity



Figure: Flaws in the mesh creating orthoradial velocity

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- The k ω has been improved to provide good agreements with experiments.
- Theoretical and Numerical analysis have been done.
- Understanding 3D results are still under investigation. Fortunately we are able to perform computations very rapidly (3 hours on 1000 processors for 10⁷ cells.

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Cea 2D Dutton experiment: data

			-1				
Aerodynamic conditions			i				
			-				
Initial wind tunnel co	onditions:						
Mach number	Mach =	2.46					
Stag. pressure	pi =	515000.00	Pa				
Stag. temperature	Ti =	293.00	K				
Reference lenght and s	rea:						
Reference lenght	Lizef =	0.25400	70				
Reference surface	Sref =	0.00317	m2				
Gas model (perfect gas	e) :						
Spec. heat ratio	Gamma =	1.400					
Ideal gas constant	rgas -	287.053	J/kg.K				
Specific heat (cst V)	Cv =	717.633	J/kg.K				
Specific heat (cst p)	Cp =	1004.686	J/kg.K				
Ideal gas cst / mole	Rgas =	8.31432	J/(mole.K)				
Holecular weight	<u>Mair</u> =	0.02896	kg/mole				
Wind tunnel conditions	**						
Temperature	т =	132.56	K				
Pressure	р -	32078.50	Pa				
Density	£9. =	0.84302	kg/m**3				
Velocity and sound speed:				Velocity and sound speed:			
verocicy magnitude	¥ =	567.79	m/ B	Velocity magnitude	v -	567.79 m/s	
found around	· -	2044.04	KIIV II	~	v =	2044.04 km/h	
sound speed	a -	230.01	inv o http://b	sound speed	a -	230.01 m/8	
	u –	030.94	Any n		a -	030.91 Km/n	
Viscosity model (sutherland law):				Viscosity model (suther	Viscosity model (sutherland law):		
Suth constant	T	110 40	v	Subb anatast	T	110 00 8	
Ref. temperature	Tref =	273.16	R .	Bef temperature	Trof =	272 16 K	
Ref. viscosity	Nuref =	0.1728-04	km/(m.s)	Def vigcosity	Huraf =	0 172E-04 km/(m m)	
	00305.865		accasoner	Neil Viboobioy	86.086.38A	or real or high lights?	
Molecular viscosity	Nu =	0.916E-05	kg/(m.s)	Molecular viscosity	M11 =	0.916E-05 kg/(m.s)	
Kinematic viscosity	Nu =	0.109E-04	m2/s	Kinematic viscosity	Nu =	0.109E-04 m2/s	
Reynolds number:				Reynolds number:			
Reynolds number / m	Re/m =	52.27 × 1	0+6	Perpolds number (m	De/m =	52 27 x 10+6	
Revnolds number / Lref	Re/L =	13.28 x 1	0+6	Reynolds number / Lref	Re/L =	13.28 × 10+6	
					and the second second		

data from Dutton(1991)

Cea 2D Dutton experiment: mesh



Figure: Icem mesh with 47888 nodes (R=32mm)

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22 Reattachment zone for 2D models



Figure: Mach number: CEA k - w, Fluent $k - \omega$ model, Herrin & Dutton experimental results, $k - \epsilon$ model in Fluent, $k - \omega$ model in Fluent)