

Hans Moosmüller

Division of Atmospheric Sciences, Desert Research Institute, Reno, NV, USA

What Determines the Temperature of the Earth?

Energy Balance: $E_{in} = E_{out}$

$$E_{in} = \pi r^2 (1 - \alpha) S$$

$$E_{out} = 4\pi r^2 \epsilon \sigma T^4$$

$$\frac{(1 - \alpha) S}{4} = \epsilon \sigma T^4$$

$$T = \sqrt[4]{\frac{(1 - \alpha) S}{4 \epsilon \sigma}} = 289 K$$

σ is the Štefan-Boltzmann constant

S is the solar irradiance ($\approx 1366 \text{ W/m}^2$)

α is the planetary albedo (≈ 0.29 due to aerosols, clouds, surface)

ϵ is the emissivity (≈ 0.61 due to greenhouse gases, clouds, surface)

Aerosol Radiative Forcing Efficiency $\Delta F_{aer}/\tau$

$$\frac{\Delta F_{aer}}{\tau} = -\frac{S_0}{2} T_{atm}^2 (1 - A_{cld}) [\bar{\beta} \omega (1 - R_{surf})^2 - 2(1 - \omega) R_{surf}]$$

Aerosol Extensive Properties

ΔF_{aer} is the aerosol forcing

τ is the aerosol layer optical depth (AOD)

$\Delta F_{aer}/\tau$ is the aerosol forcing efficiency

Aerosol Intensive Properties

$\bar{\beta}$ is the average upscatter fraction

ω is the aerosol single scattering albedo (SSA)

Surface Properties

R_{surf} is the surface albedo

Atmospheric Properties

T_{atm} is the transmittance of the atmosphere above

A_{cld} is the cloud fraction

Solar Properties

S_0 is the solar constant

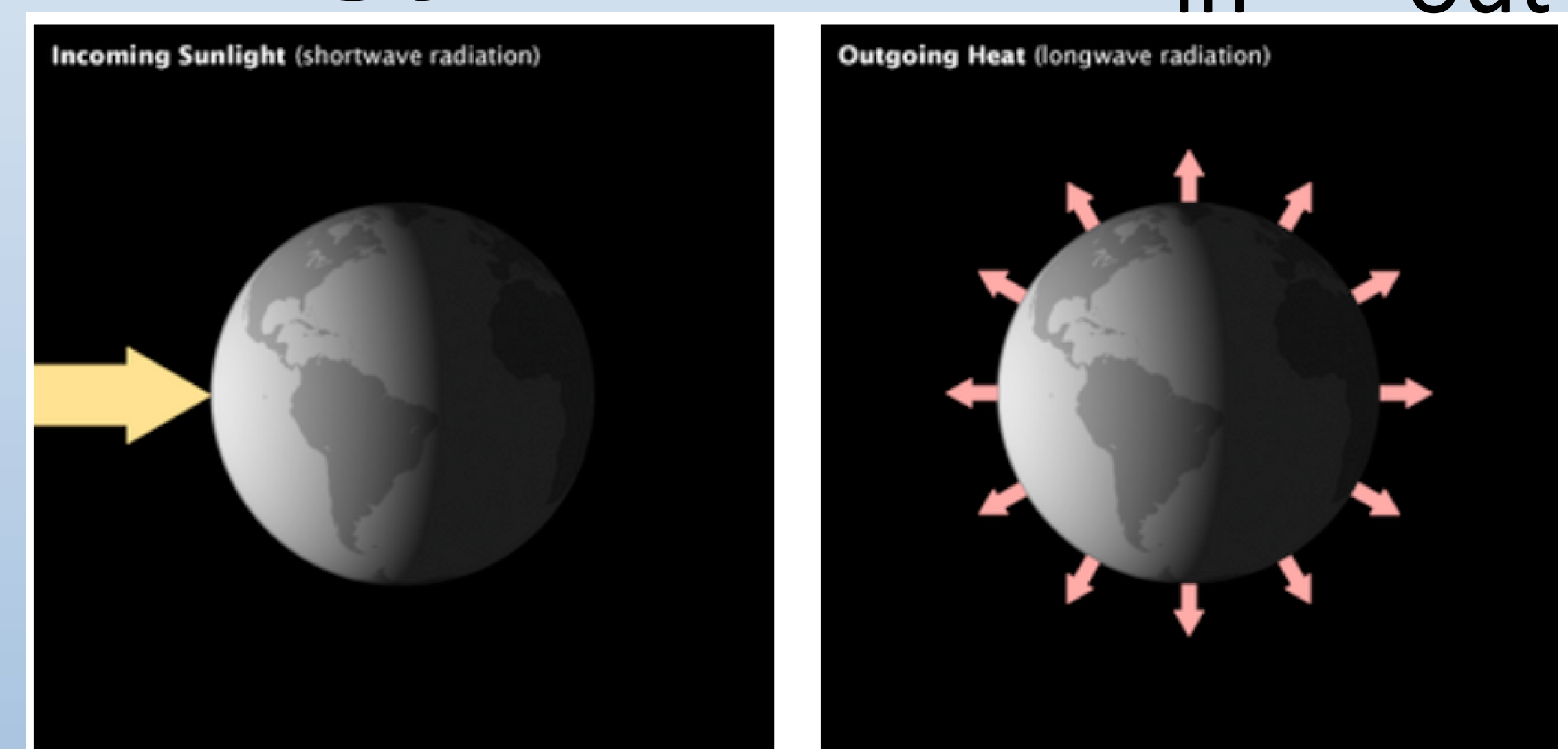
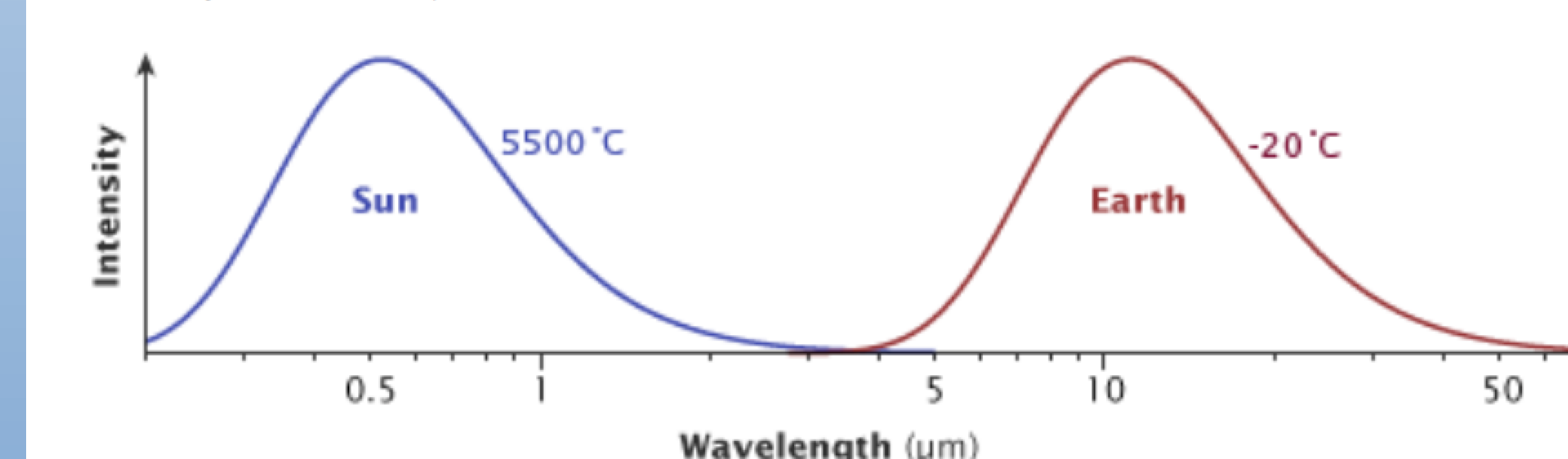


Figure 1. The energy that Earth receives from sunlight is balanced by an equal amount of energy radiating into space. The energy escapes in the form of thermal infrared radiation: like the energy you feel radiating from a heat lamp. (NASA illustrations by Robert Simmon.)



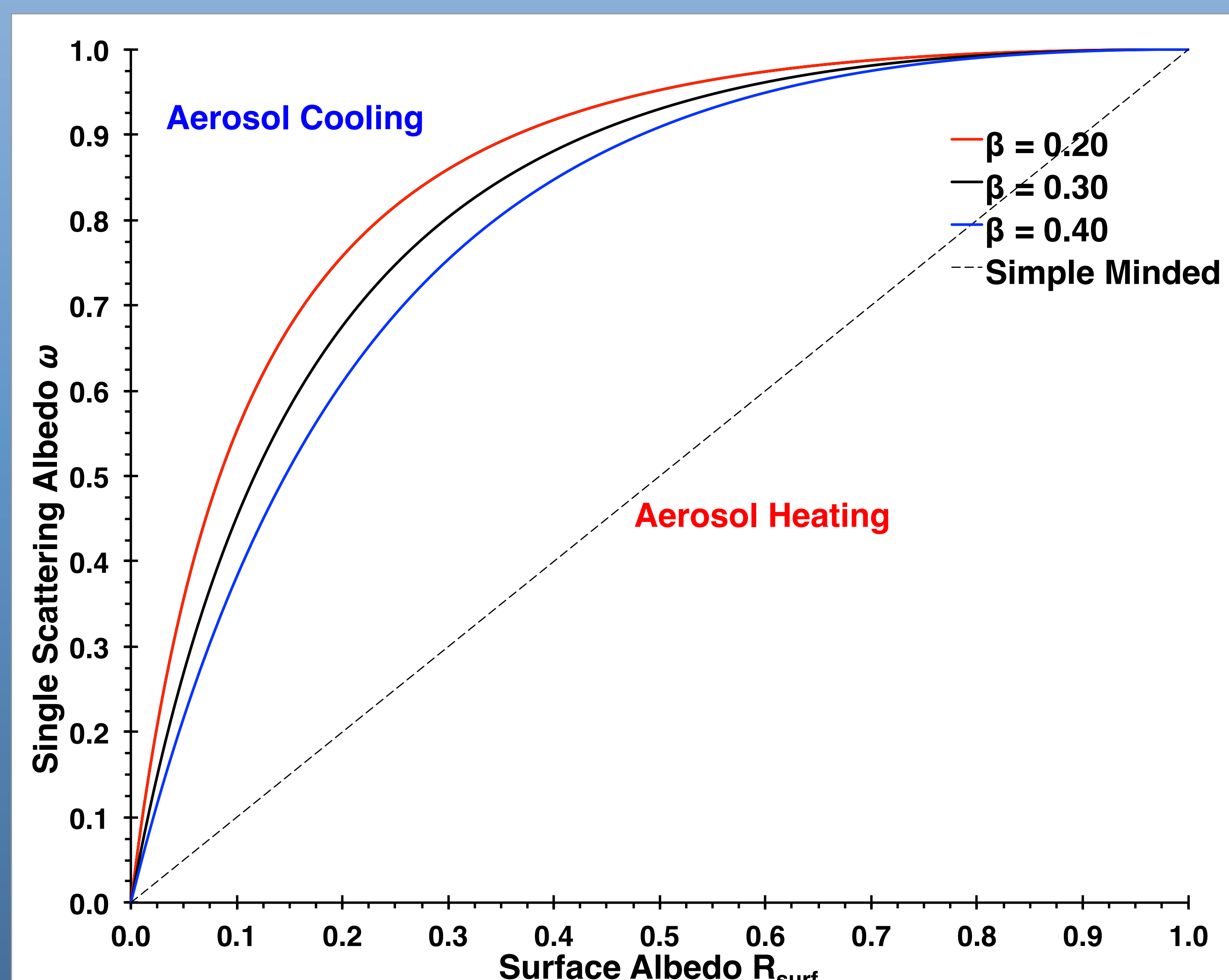
Heating or Cooling?

$\Delta F_{aer} = 0$ yields the boundary between Cooling and Heating depends only on ω , R_{surf} , and β

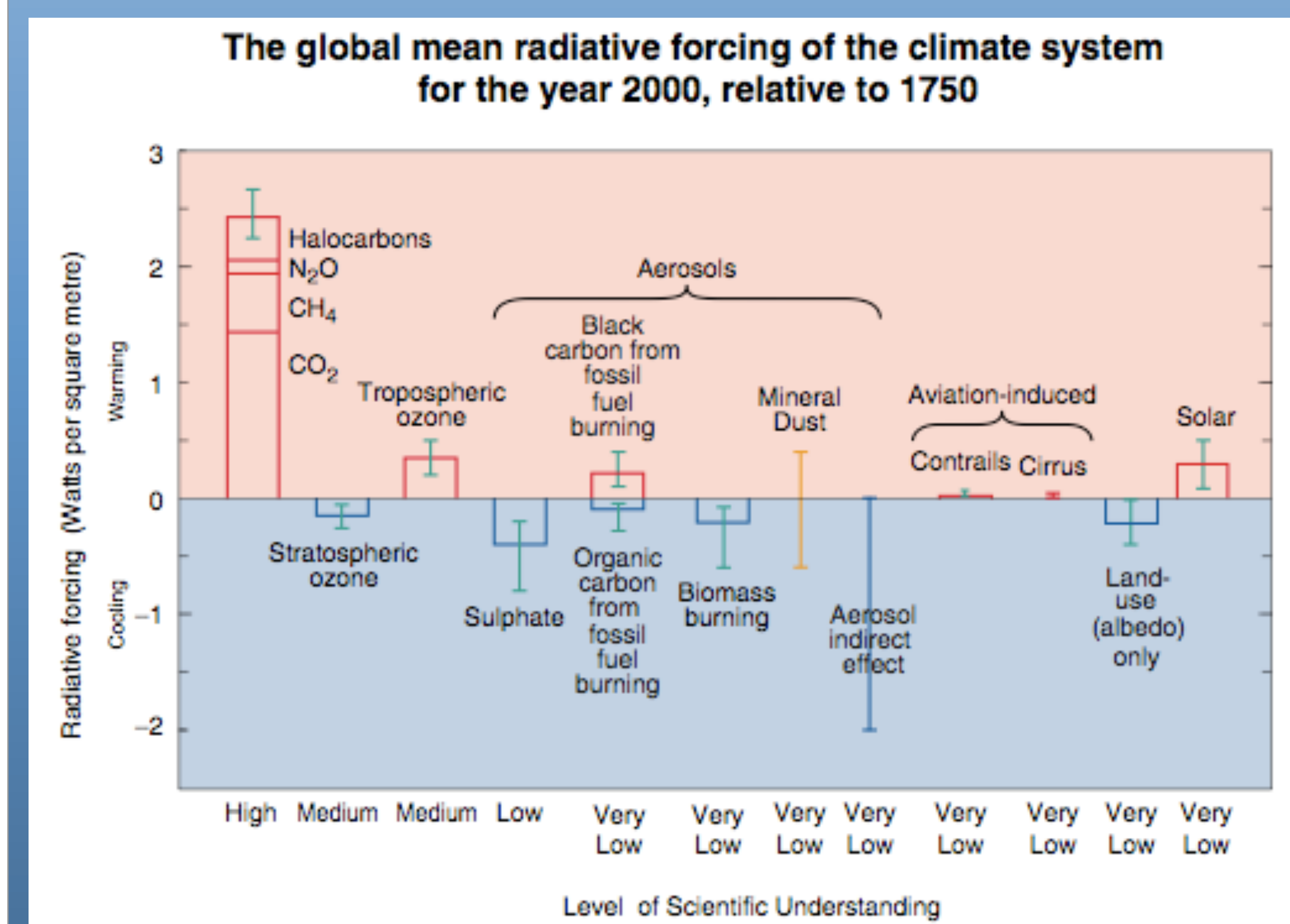
Atmospherically relevant upscatter fraction i.e., $0.2 < \beta < 0.4$ (Sheridan & Ogren; 1999)

Assumption of Henyey-Greenstein phase function

“Simple Minded” doesn’t take multiple interactions into account



Importance of Aerosol Forcing



What Determines Albedo?

1) Material Properties

Complex Refractive Index

$$m = n + ik$$

For BC: $n = 1.95$ $k = 0.79$

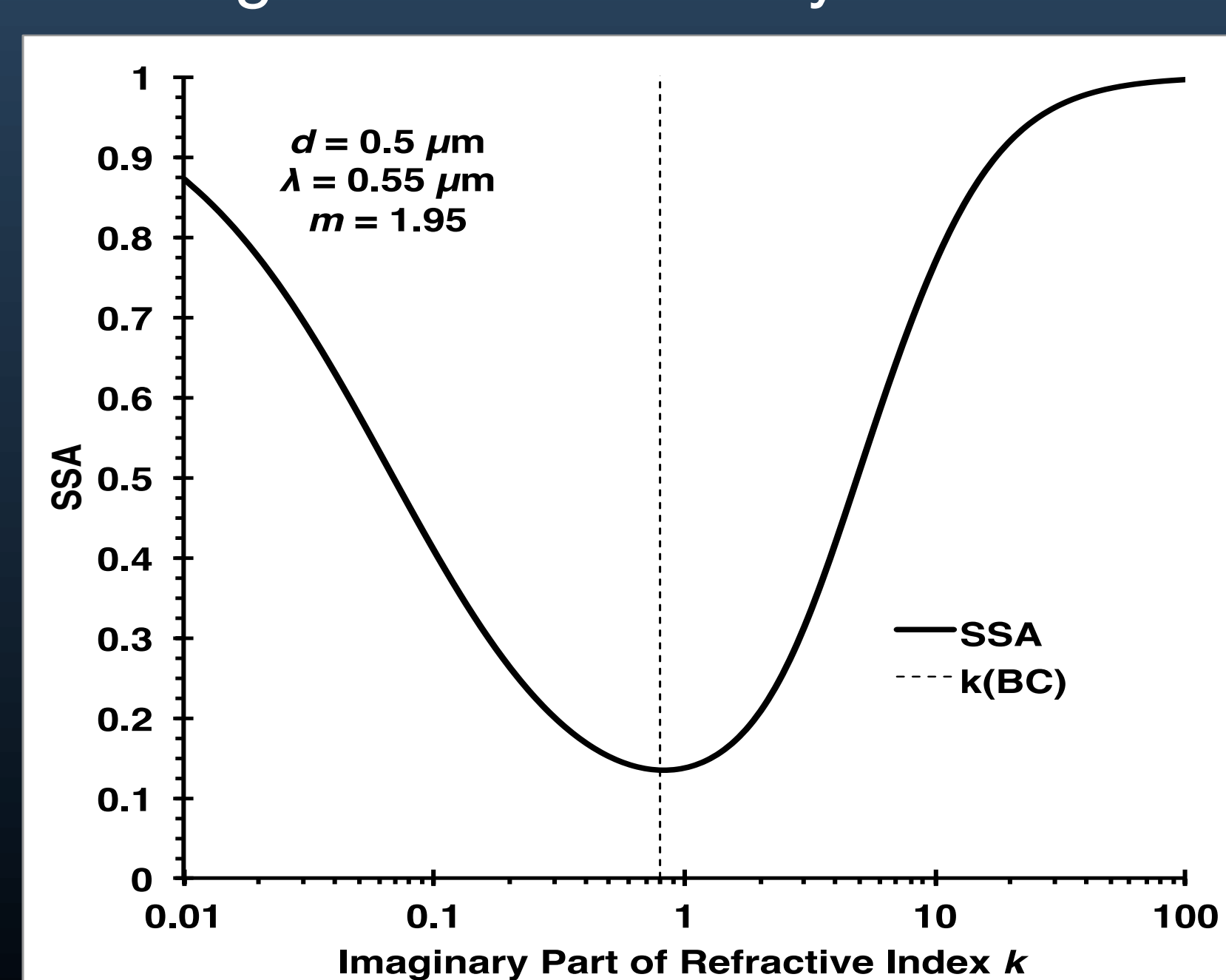
Independent of wavelength in the visible

$K = 0 \Rightarrow$ no absorption

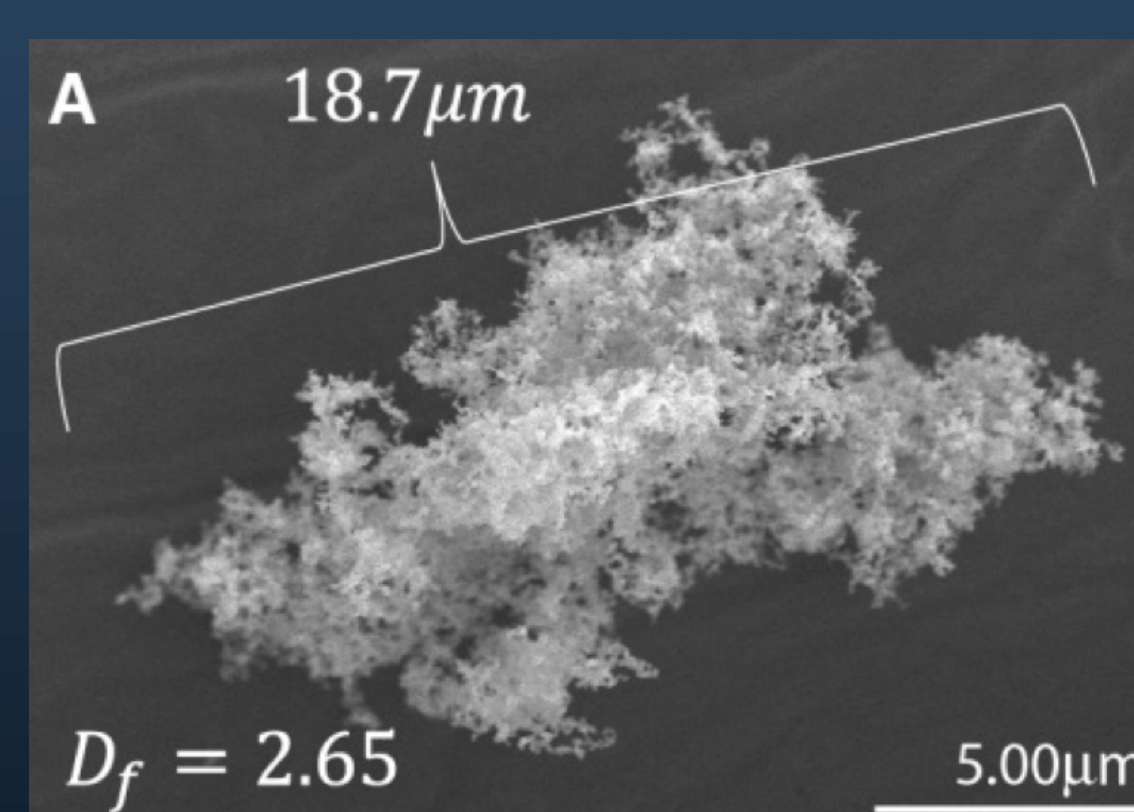
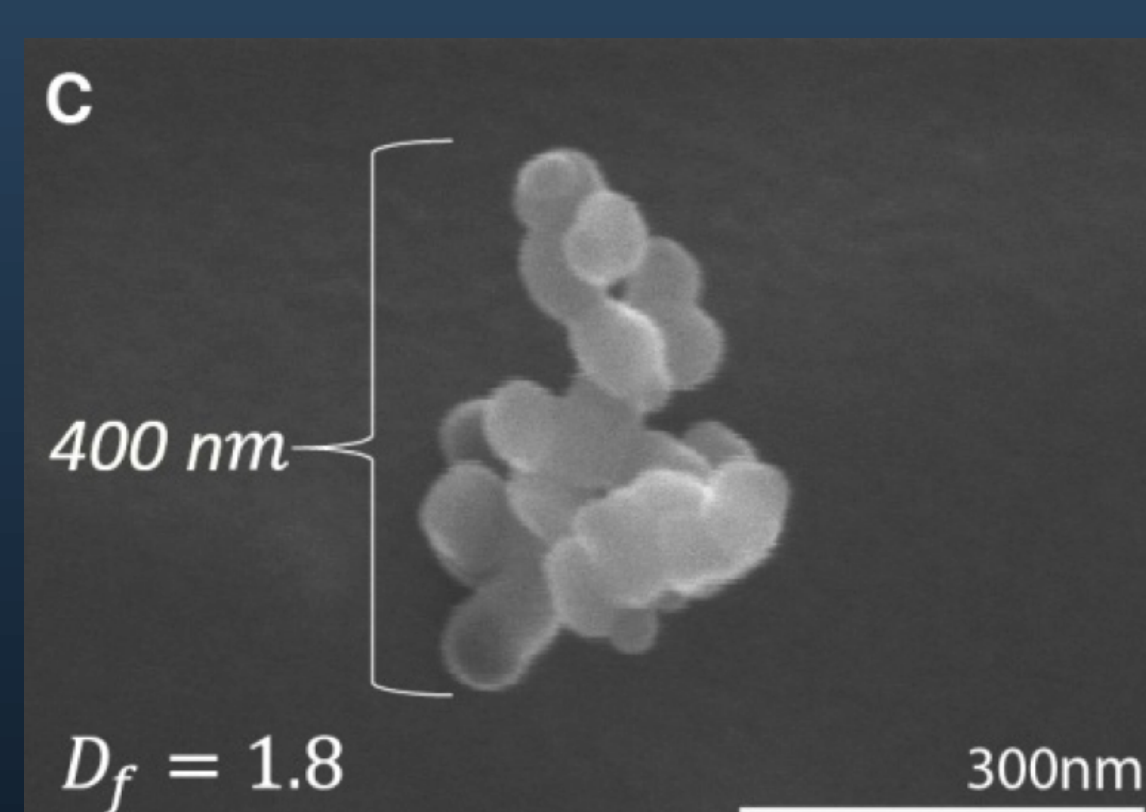
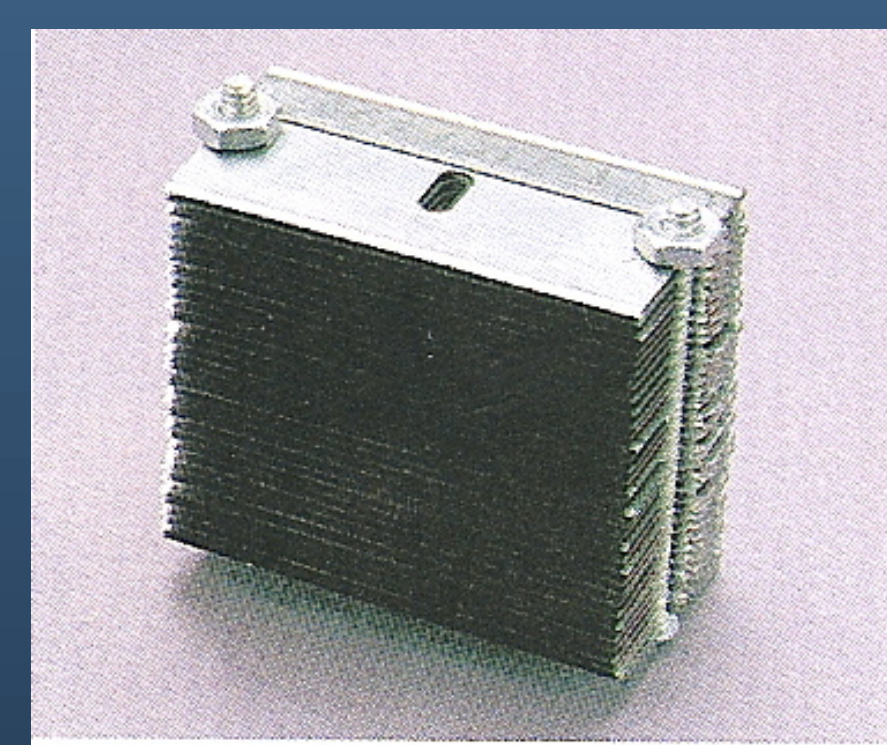
Bulk Absorption Coefficient $a_{abs} = 4\pi k/\lambda$

Penetration Depth $1/a_{abs} \sim 55 \text{ nm}$

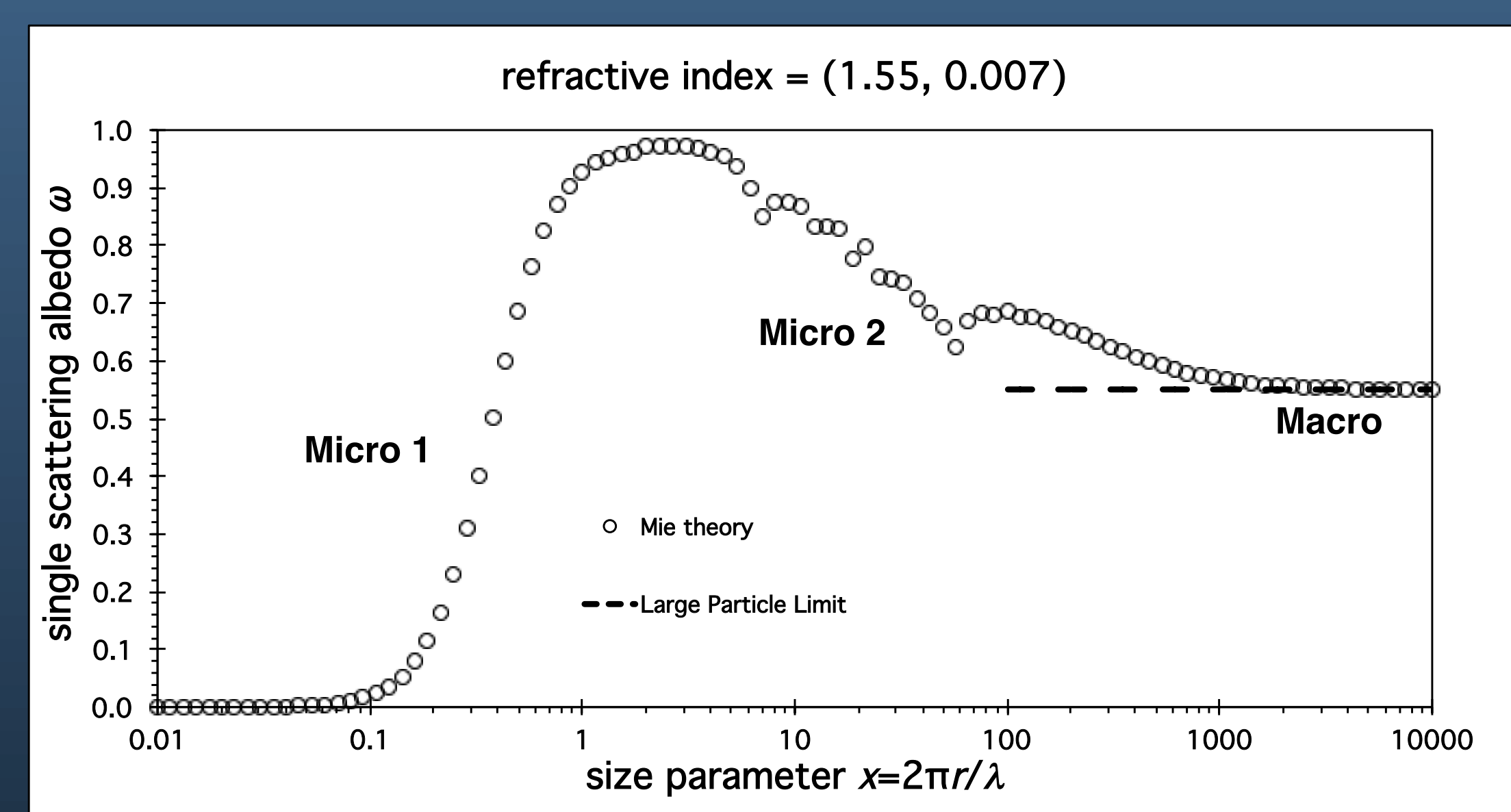
But larger k doesn’t always reduce SSA



2) Morphology



3) Size



Macroscopic: Albedo Independent of Size

Microscopic 1: Albedo Increases with Size

Microscopic 2: Albedo Decreases with Size